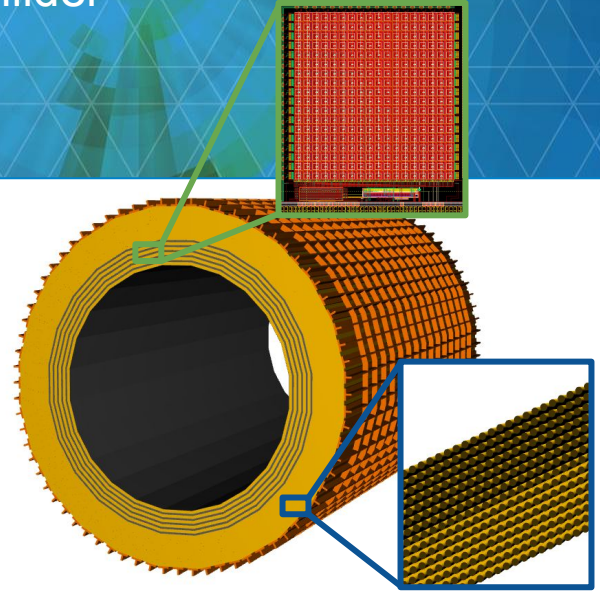


# APCTP 2nd Workshop on the Physics of Electron Ion Collider

November 29, 202

## The ePIC *Barrel Imaging Calorimeter (BIC)*



Sylvester Joosten  
Argonne National Laboratory

# Why electromagnetic calorimetry at EIC is hard

## From the EIC Yellow Report: stringent barrel ECal requirements

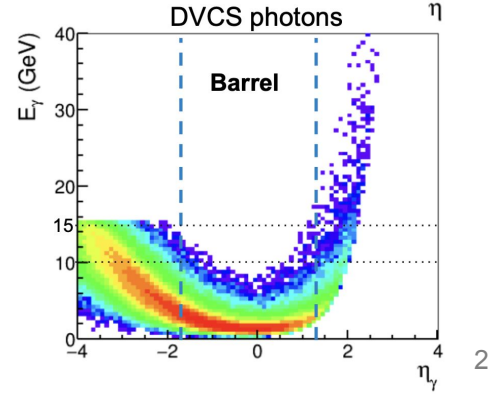
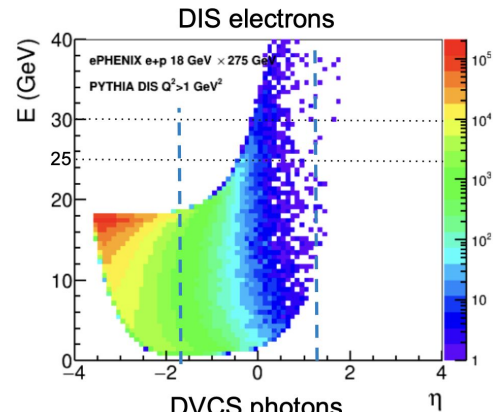
EIC is an **electron scattering** machine and identifying scattered electrons mainly depends on the electromagnetic calorimetry.

The electromagnetic calorimeter is the main detector for **electron-pion separation**. The inclusive physics program requires up to  $10^4$  pion suppression at low momenta in the barrel.

The exclusive program requires **decent energy resolution** ( $< 7\%/\sqrt{E} \oplus 1-3\%$ ) for **photon energy reconstruction**, and also the **fine granularity** for good  $\pi^0$ - $\gamma$  separation up to 10 GeV.

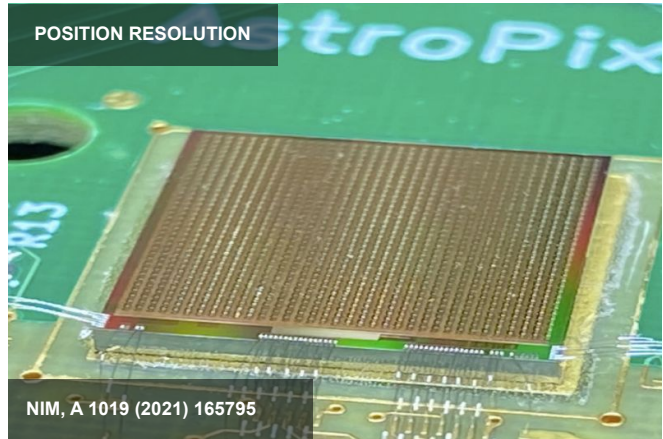
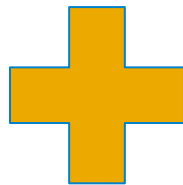
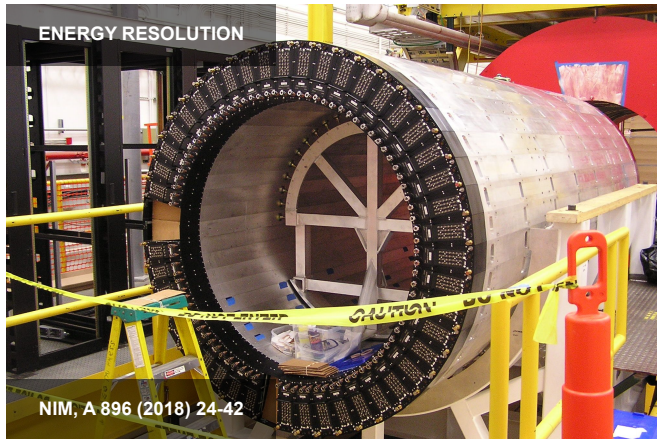
The bECal should be capable of measuring **low energy photons** down to 100 MeV, while having the range to measure energies well above 10 GeV

The system is space-constrained to very **limited space** inside the solenoid.



# CONCEPT: A HYBRID IMAGING CALORIMETER

Combination of a high-performance sampling calorimeter with inexpensive silicon sensors for shower profiling

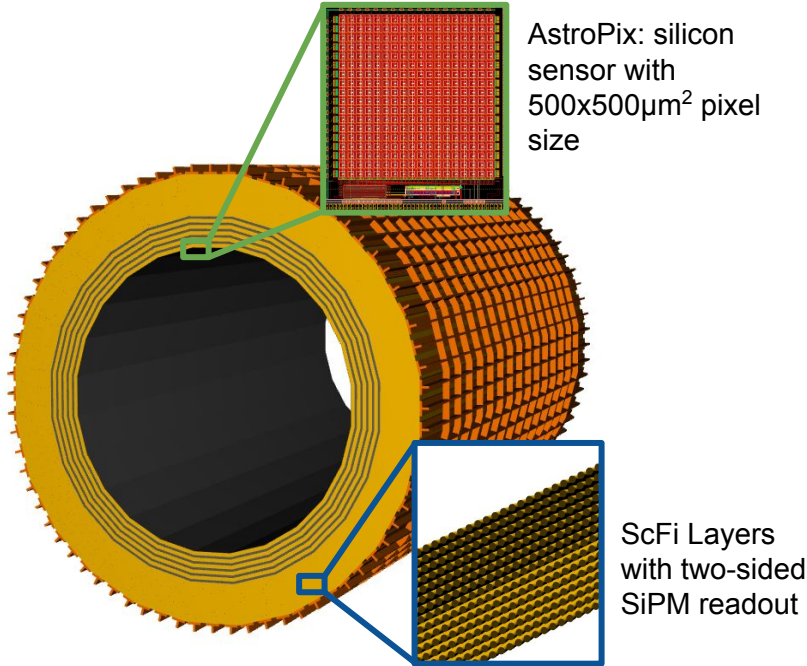


Start from mature layered Pb/ScFi technology with side-readout (same as the GlueX calorimeter) for state-of-the-art sampling calorimeter performance

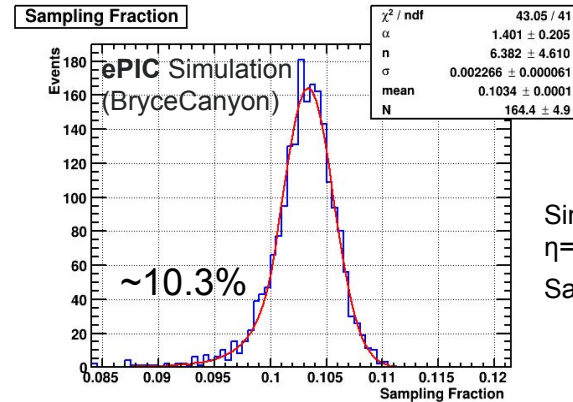
Insert layers of monolithic AstroPix sensors (inexpensive ultra-low-power silicon sensor developed for NASA) in the first half of the calorimeter to capture a 3-D image of the developing shower

# General Overview

## BARREL IMAGING CALORIMETER (BIC)



- **4(+2) layers of imaging Si sensors interleaved with 5 Pb/ScFi layers**
- Followed by a **large section of Pb/ScFi section**
- Total radiation thickness  $> 17.1 X_0$
- Sampling fraction  $\sim 10\%$



Simulation of **single photons** at  $\eta=0$  ( $\sim 17.1 X_0$ )

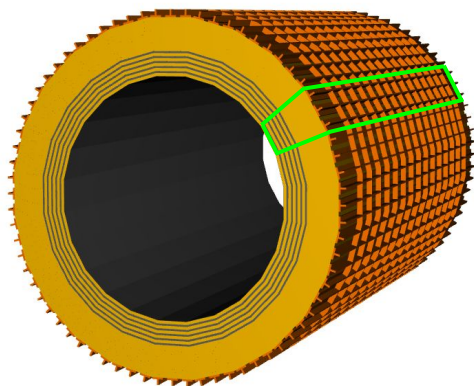
Sampling fraction =  $\frac{\Sigma E_{\text{fibers}}}{E_{\text{thrown}}}$

**Energy resolution** - Primarily from Pb/ScFi layers (+ Imaging pixels energy information)

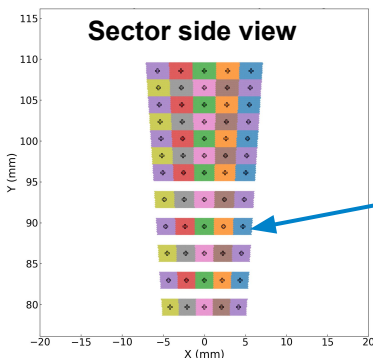
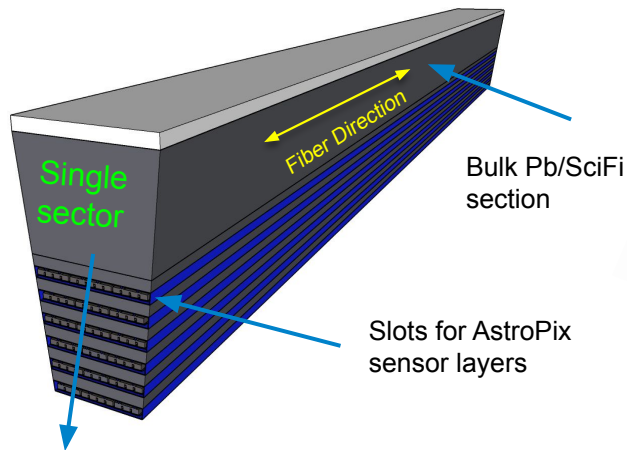
**Position resolution** - Primarily from Imaging Layers (+ 2-side Pb/ScFi readout)

# Detector Structure

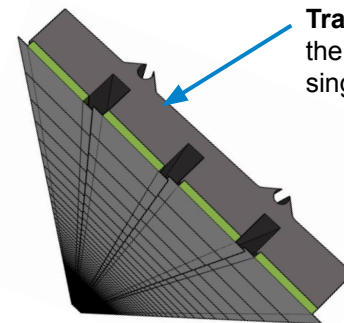
## BARREL IMAGING CALORIMETER (BIC)



**Length:** 432.5 cm  
**Radius:** ~ 80 cm radius,  
**Structure:** 48 sectors  
 **$\eta$  Range:**  $-1.71 < \eta < 1.31$



**Pb/SciFi Layer** - 12 layers per sector  
**Structure:** 5 readout cells (one light-guide per readout cell)  
**Construction:** 17 rows of fiber



**Tray** - Structure holding the AstroPix staves for a single layer

**Length:** ~ 200 cm (half length)  
**Structure:** 6-7 “turbofanned” staves per tray  
**Stave Structure:** ~ 13 Modules per stave

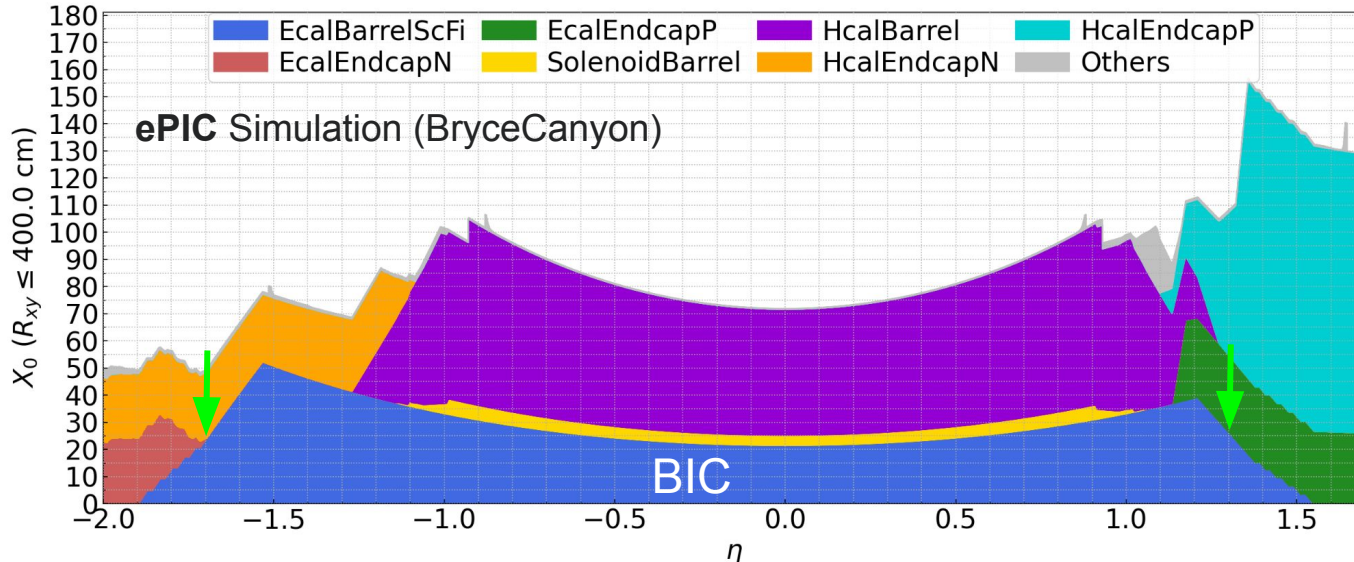


**Length:** ~ 16 cm  
**Width:** ~ 2 cm  
**Gaps:** < 200  $\mu$ m  
**Structure:** ~ 8 chips/module

**Module** - Several AstroPix chips daisy-chained together on Flex PCB

# CALORIMETER COVERAGE AT EPIC

- Non-projective “continuous” design
- Particles passing at steep angle pass through *much* more material than at central rapidities (up to 45 X0 at negative  $\eta$ )
- BIC is responsible for  $-1.7 < \eta < 1.3$
- Continuous transition between backward-barrel-forward calorimeters



# PB/SCFI TECHNOLOGY

## Our Pb/ScFi layers follow the GlueX Design

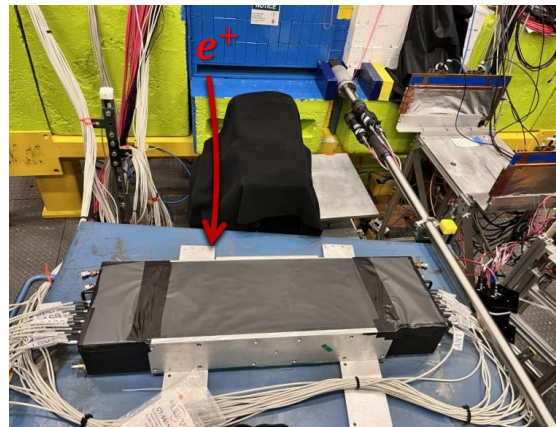
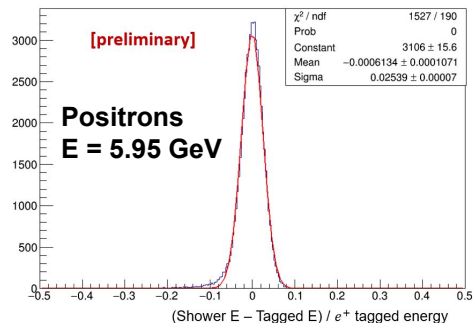
- Mature Technology: **GlueX**, KLOE electromagnetic calorimeters
  - Detailed studies on **calorimetry performance**, including the light collection uniformity in fibers, light collection efficiencies, etc.
  - **Module construction** (lead handling, swaging, Pb/ScFi layers assembly, module machining) fully developed for GlueX  
*Z. Papandreou, <https://halldweb.jlab.org/DocDB/0031/003164/>*
  - Assembly and installation of self-supporting barrel based on GlueX experience
- Tested extensively for electromagnetic response in energies  $E_{\gamma} < 2.5 \text{ GeV}$
- **Energy resolution:**  $\sigma = 5.2\% / \sqrt{E} \oplus 3.6\%^{(1)}$ 
  - $15.5 X_0$ , GlueX could not constrain the constant term due to low energies
  - New results from Hall D beam tests show that constant term  $< 2\%$

### March 2023 beam test

**Measured Resolution:  $\sim 2.5\%$**

Extrapolated GlueX NIM<sup>1)</sup>:  $\sim 4.2\%$

**Trends well below a 2% constant term!**



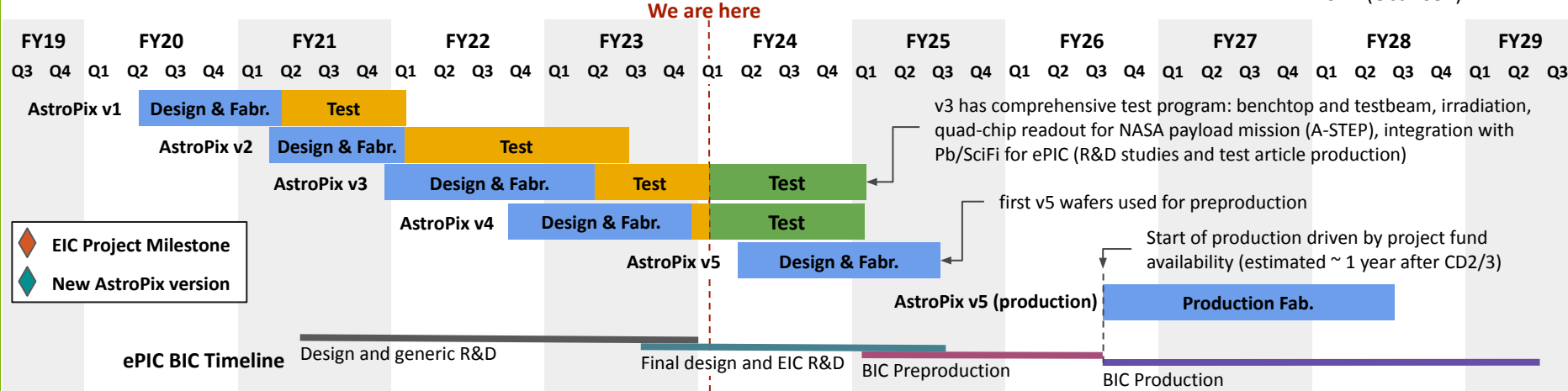
Baby BCAL  
60 cm long  
15.5 X<sub>0</sub>

tested with  $e^+$   
E  $\sim 3.6\text{-}6 \text{ GeV}$

1) GlueX, Nucl. Instrum. Meth. A, vol. 896, pp. 24–42, 2018

# ASTROPIX TECHNOLOGY

**Not shown:**  
 Early CD4 (Oct 2032)  
 CD4 (Oct 2034)



EIC Project Milestone  
 New AstroPix version

v3 has comprehensive test program: benchtop and testbeam, irradiation, quad-chip readout for NASA payload mission (A-STEP), integration with Pb/SciFi for ePIC (R&D studies and test article production)

first v5 wafers used for preproduction

Start of production driven by project fund availability (estimated ~ 1 year after CD2/3)

**ePIC BIC Timeline**

Design and generic R&D      Final design and EIC R&D      BIC Preproduction      BIC Production

CD0      CD1      CD3a      CD3b      CD2/3      Start of BIC installation at BNL



**AstroPix v1**  
 HV-CMOS MAPS based on ATLASPix3, designed for the AMEGO-X NASA mission, optimized for power dissipation and energy resolution  
*Nucl.Instrum.Meth.A 1019 (2021) 165795*

0.45 x 0.45 cm<sup>2</sup> chip, 175 μm pixel pitch  
 18 x 18 pixel matrix  
 Power dissipation 14.7 mW/cm<sup>2</sup>



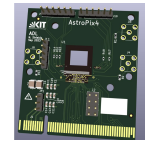
**AstroPix v2**

1 x 1 cm<sup>2</sup> chip, 250 μm pixel pitch  
 35 x 35 pixel matrix  
 Row/column readout  
 Power dissipation 3.4 mW/cm<sup>2</sup>



**AstroPix v3**

2 x 2 cm<sup>2</sup> chip, 500 μm pixel pitch  
 Row/column readout  
 Power dissipation <1 mW/cm<sup>2</sup>  
 2.5 MHz timestamp, 200 MHz ToT



**AstroPix v4**

1 x 1 cm<sup>2</sup> chip, 500 μm pixel pitch  
 Individual pixel readout  
 3 timestamps, 3.25ns time resolution  
 TuneDAC for pixel-by-pixel thresholds

**AstroPix v5**

2 x 2 cm<sup>2</sup> chip, 500 μm pixel pitch  
 Design identical to v4 (with bug fixes)



# ASTROPIX LAYER SIZE

The BIC will be a large silicon detector:

- 4 (+2) layers in the ePIC barrel will cover  $-1.7 < |\eta| < 1.3$
- The Astropix sensor area will be about  $100 \text{ m}^2$  (+40  $\text{m}^2$ )
- ~ 250,000 (+125,000) chips

## Other comparable Si detector arrays in advanced stage

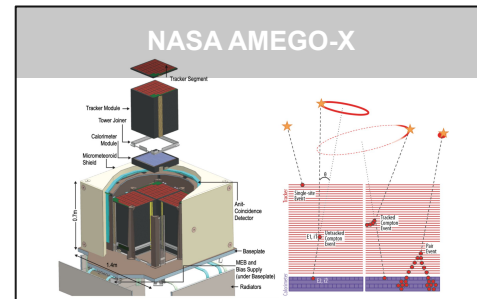
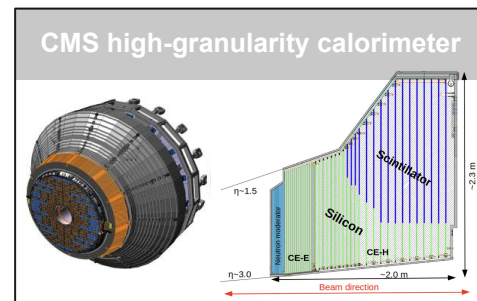
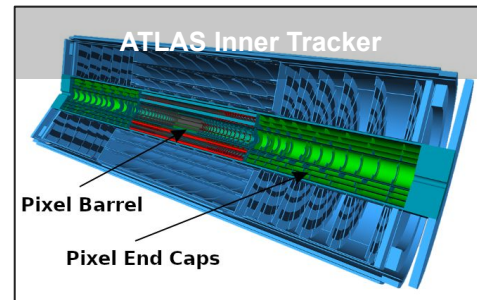
- ATLAS Inner Tracker - silicon strips<sup>1</sup> (ITk pixel)  $160 \text{ m}^2$  (50 million channels)
- CMS high granularity calorimeter <sup>2</sup> ~  $600 \text{ m}^2$  (6.5 million channels)
- **AMEGO-X NASA mission:**
  - Will use a  $40 \text{ m}^2$  AstroPix-based tracker, to be sent into space
  - We plan to use chips off-the-shelf: **no design modifications.**

## Advantages of AstroPix with respect to pixels used in e.g. ATLAS

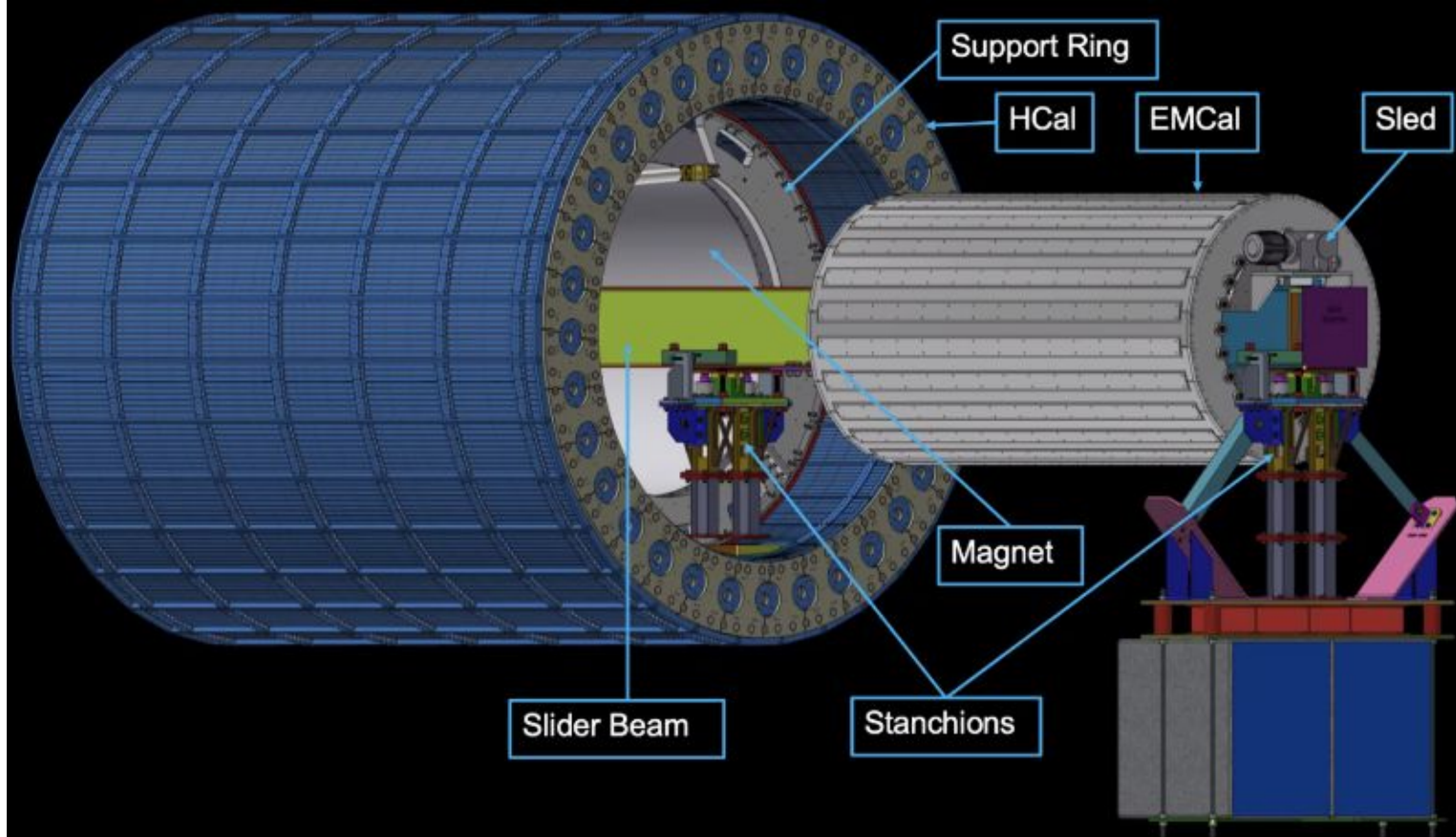
- AstroPix has very low power consumption (used in space)
  - 100 times smaller power consumption per  $\text{cm}^2$  than ATLASPix pixels
  - AstroPix is a monolithic sensor - less complicated structure
  - No bump bonding - less risk of damaging sensors

<sup>1</sup> arXiv:2105.10367, ATLAS ITk Pixel Detector Overview

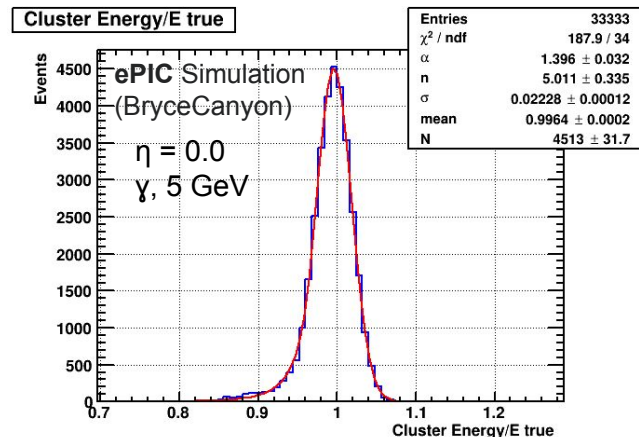
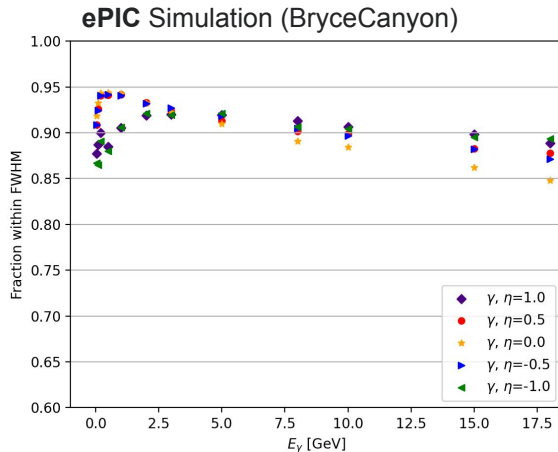
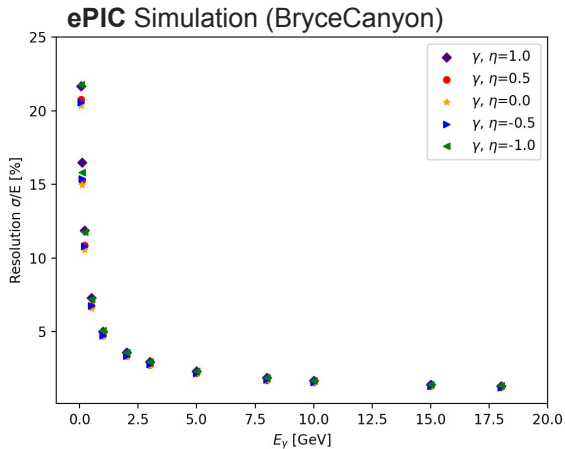
<sup>2</sup> arXiv:1802.05987, The CMS High-Granularity Calorimeter for Operation at the High-Luminosity LHC2



# ASSEMBLY TOOLING



# ENERGY RESOLUTION - PHOTONS



## Fit parameters

$\eta$	$a/\sqrt{E}$ [%]	$b$ [%]
-1	5.1(0.01)	0.47(0.03)
-0.5	4.77(0.01)	0.38(0.02)
0	4.67(0.01)	0.40(0.02)
0.5	4.75(0.01)	0.39(0.02)
1	5.1(0.01)	0.41(0.02)

- Based of Pb/ScFi part of the calorimeter
- Resolution extracted from a Crystal Ball fit  $\sigma$

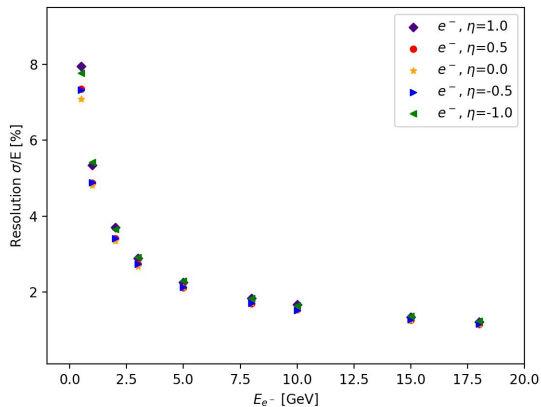
**GlueX Pb/ScFi ECal:**  $\sigma = 5.2\% / \sqrt{E} \oplus 3.6\%$  NIM, A 896 (2018) 24-42

- $15.5 X_0$ , extracted for integrated range over the angular distributions for  $\pi^0$  and  $\eta$  production at GlueX ( $E_\gamma = 0.5 - 2.5 \text{ GeV}$ )
- Measured energies not able to fully constrain the constant term

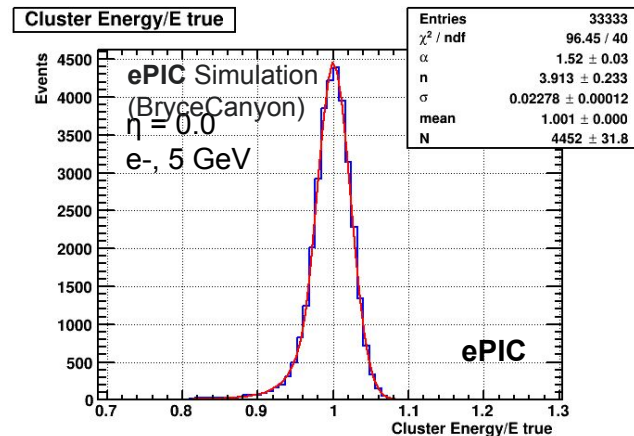
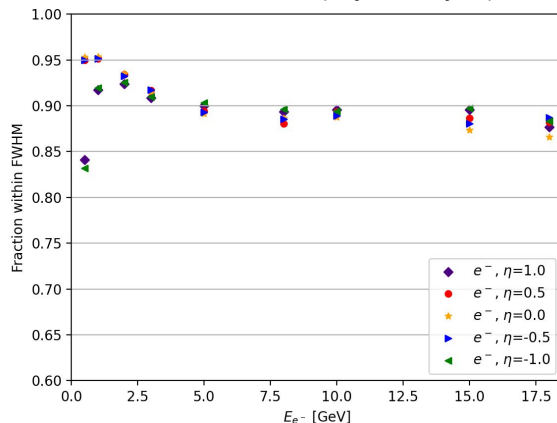
Simulations of **GlueX prototype** in ePIC environment agree with data at  $E_\gamma < 0.5 \text{ NIM}$ , 596 (2008) 327-337

# ENERGY RESOLUTION - ELECTRONS

ePIC Simulation (BryceCanyon)



ePIC Simulation (BryceCanyon)



Fit parameters

$\eta$	$a/\sqrt{E}$ [%]	$b$ [%]
-1	5.22(0.02)	0(0.08)
-0.5	4.88(0.01)	0(0.04)
0	4.81(0.01)	0(0.08)
0.5	4.88(0.01)	0(0.04)
1	5.19(0.01)	0(0.06)

Resolution extracted from a crystal ball fit  $\sigma$

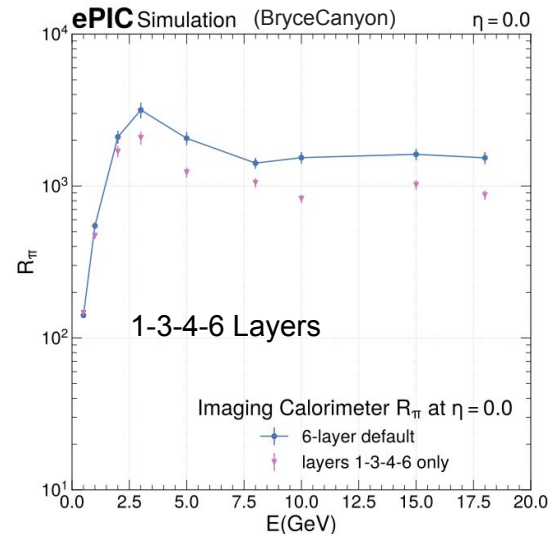
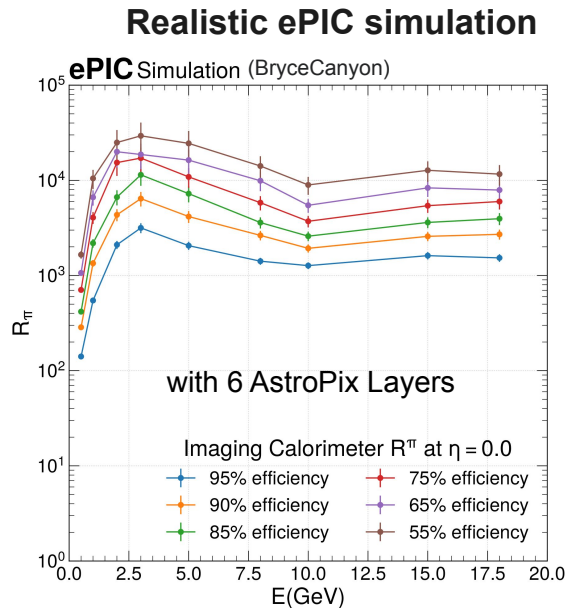
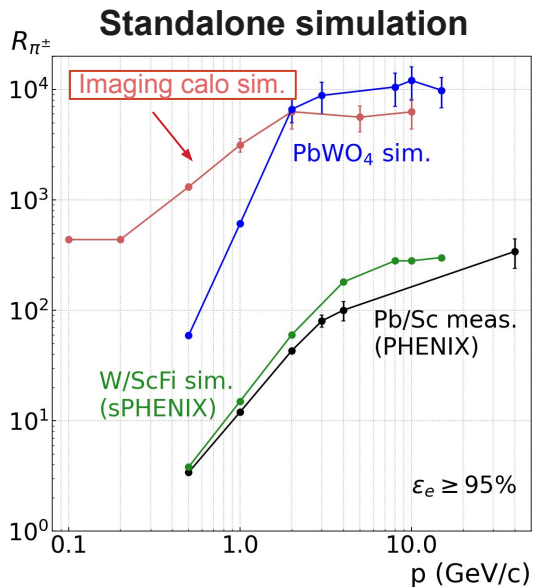
**GlueX Pb/ScFi ECal:**  $\sigma = 5.2\% / \sqrt{E} \oplus 3.6\%$  NIM, A 896 (2018) 24-42

- $15.5 X_0$ , extracted for integrated range over the angular distributions for  $\pi^0$  and  $\eta$  production at GlueX ( $E_\gamma = 0.5 - 2.5$  GeV)
- Measured energies not able to fully constrain the constant term

Simulations of **GlueX prototype** in ePIC environment agree with data at  $E_\gamma < 0.5$  NIM, 596 (2008) 327-337

# BIC: PERFORMANCE EXAMPLE

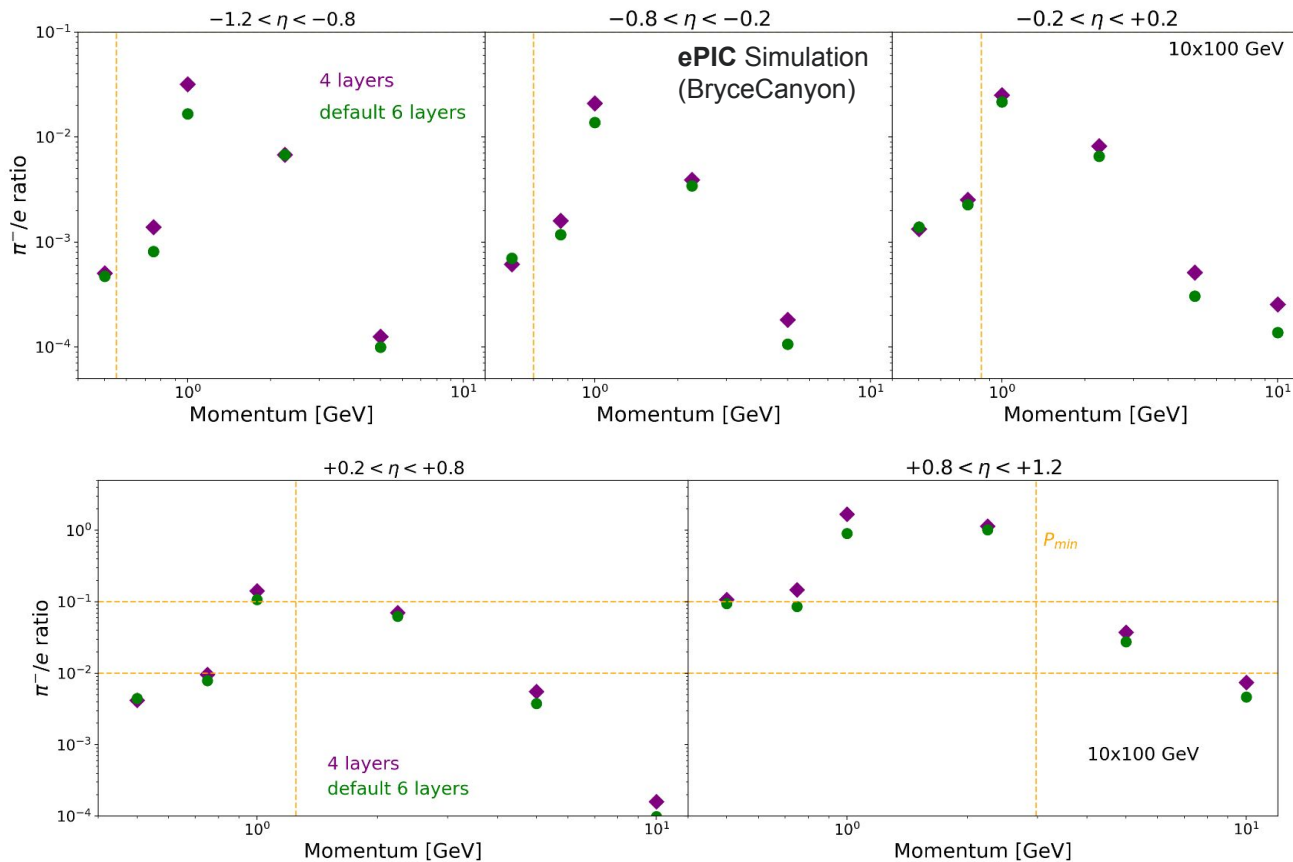
## e/ $\pi$ Separation



- **Goal:** Separation of electrons from background in Deep Inelastic Scattering (DIS) processes
- Method: **E/p cut (Pb/ScFi) + Neural Network** using **3D position and energy info** from imaging layers
- e- $\pi$  separation exceeds  $10^3$  in pion suppression at **95% efficiency** above 1 GeV in realistic conditions!

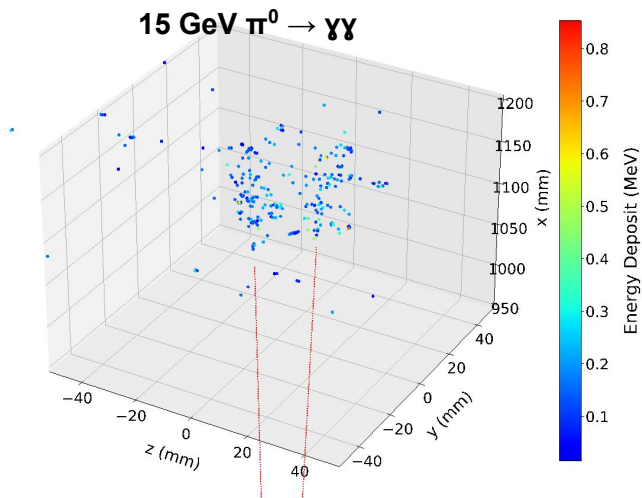
# BIC: PERFORMANCE EXAMPLE

## Impact on Inclusive Physics

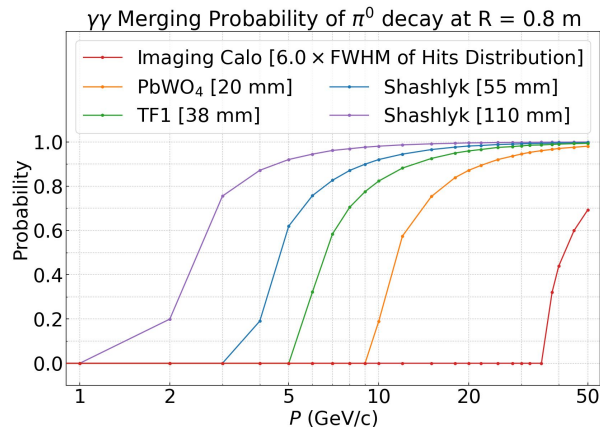


10 x 100 GeV

# NEUTRAL PION IDENTIFICATION

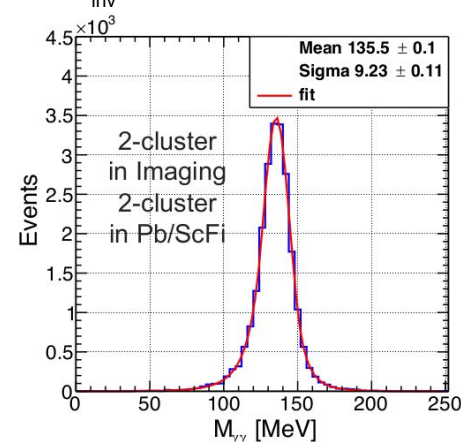


## Separation of $\gamma/\pi^0$ (upper limit)



ePIC Simulation (BryceCanyon)

$M_{inv}$  reconstruction 2 GeV  $\pi^0$



- **Goal:** Discriminate between  $\pi^0$  decays and single  $\gamma$  from DVCS, neutral pion identification
- Precise position resolution allow for excellent separation of  $\gamma/\pi^0$  based on the 3D shower profile
- Reconstruction of 2 GeV  $\pi^0$  invariant mass as a testing ground for cluster energy splitting

Separation of two gammas from neutral pion well above required 10 GeV

# EXPLORATORY: $\gamma/\pi^0$ SEPARATION

**Convolutional neural network** utilizing energy and spatial information from AstroPix layers

- Started from **10 GeV/c at  $\eta = 0$**  - the upper limit for  $\gamma/\pi^0$  from YR

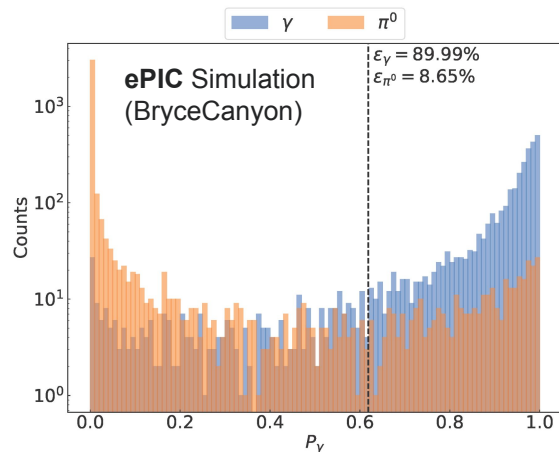
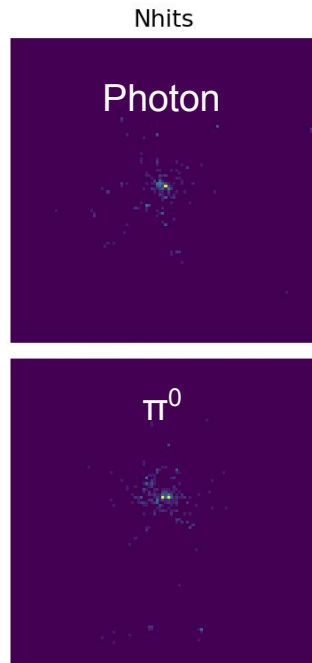
No proper **topological clustering algorithm** in the ePIC reconstruction yet

**With a quick study we easily achieved**

10 GeV/c particles - **91.4%** rejection of  $\pi^0$  at **90%** efficiency of  $\gamma$  (better than  $\text{PbWO}_4$  crystal with 20mm block size)

**Full study is ongoing:**

- Implementing optimized topological clustering for AstroPix layers
- Significant improvements expected





USA

Argonne National Laboratory



NASA Goddard Space Flight Center



Oklahoma State University



University of Connecticut



University of California Santa Cruz



Canada

University of Manitoba



University of Regina



Mount Allison University



NSERC



Canada Fund for Innovation



Korea

Kyungpook National University



Yonsei University



University of Seoul



Pusan National University



Korea University



Sungkyunkwan University



Hanyang University



Gangneung-Wonju National University



Germany

Karlsruhe Institute of Technology

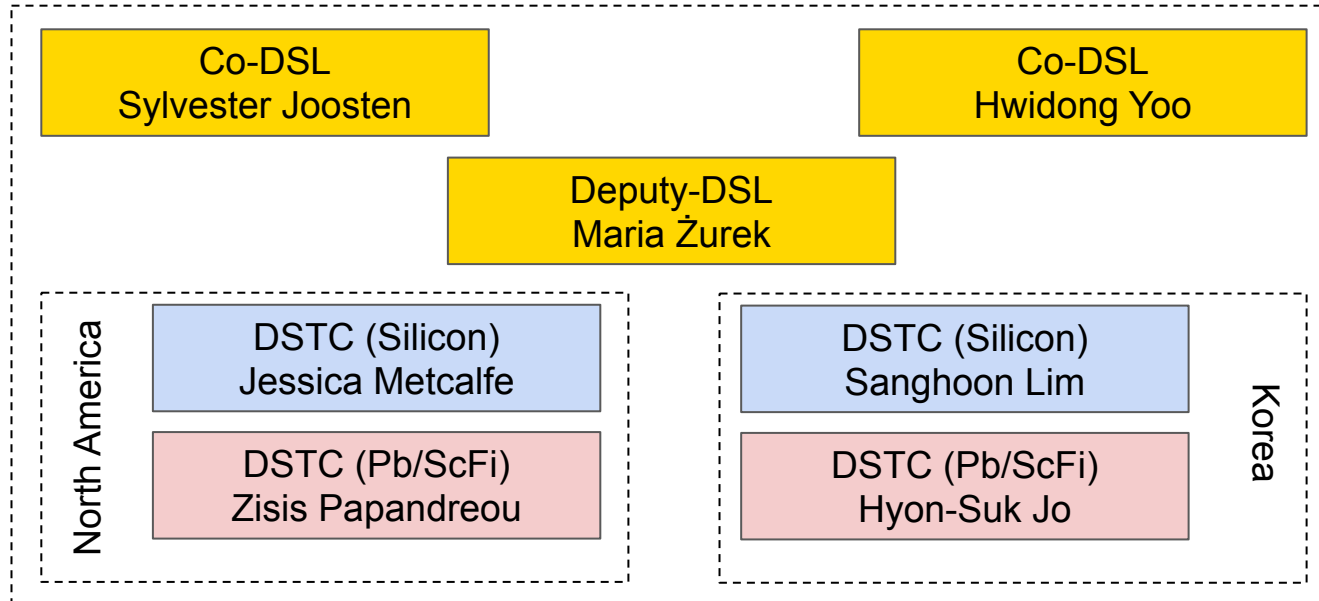


University of Giessen

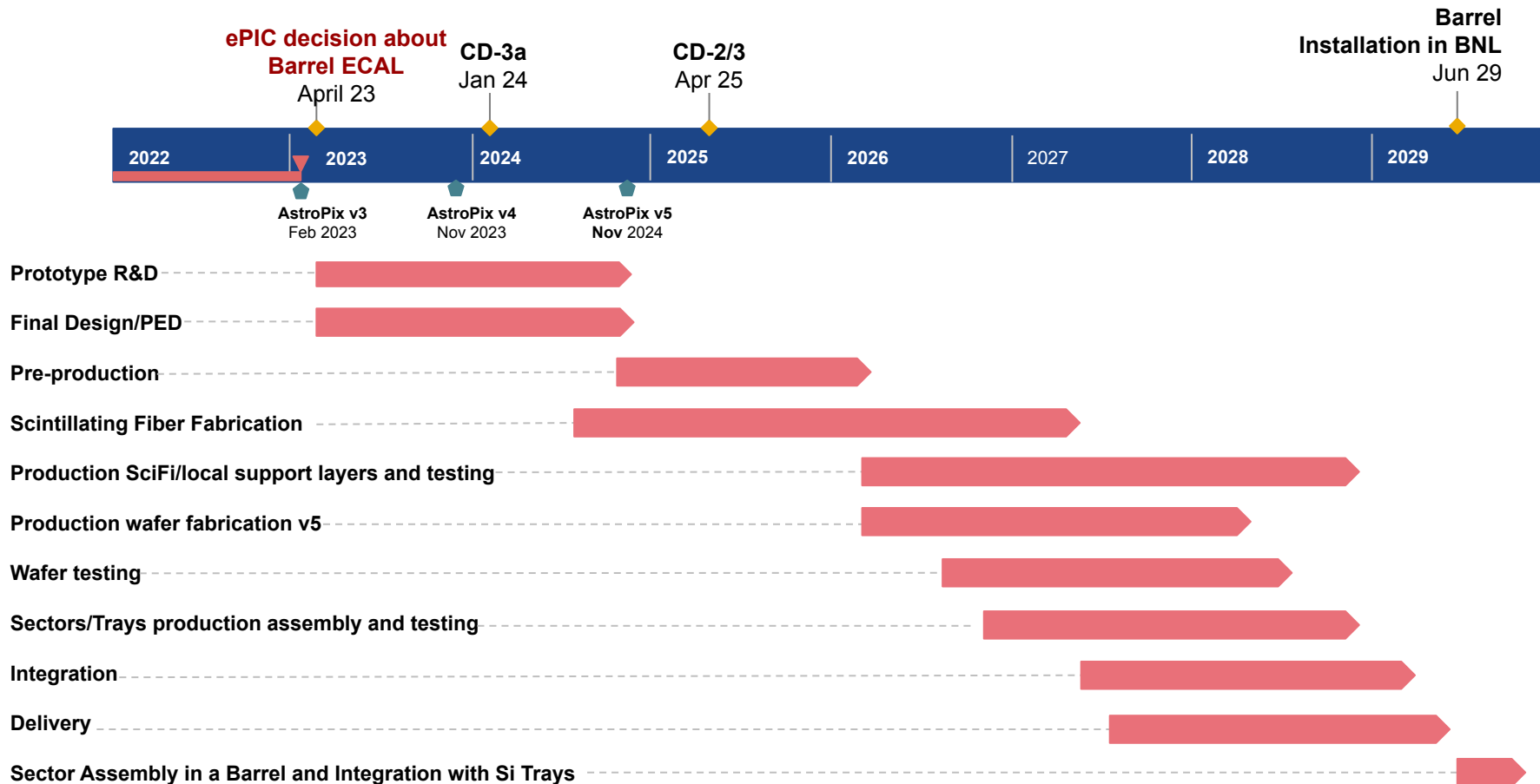


# ePIC BIC Detector Subsystem Collaboration

# BIC DETECTOR SUBSYSTEM COLLABORATION



# BIC HIGH-LEVEL SCHEDULE



# OPEN R&D QUESTIONS

To be completed with the R&D program before CD-3

**How detector performance obtained from detailed simulations compare with the measurements in the integrated SciFi/Pb and AstroPix prototype system?**

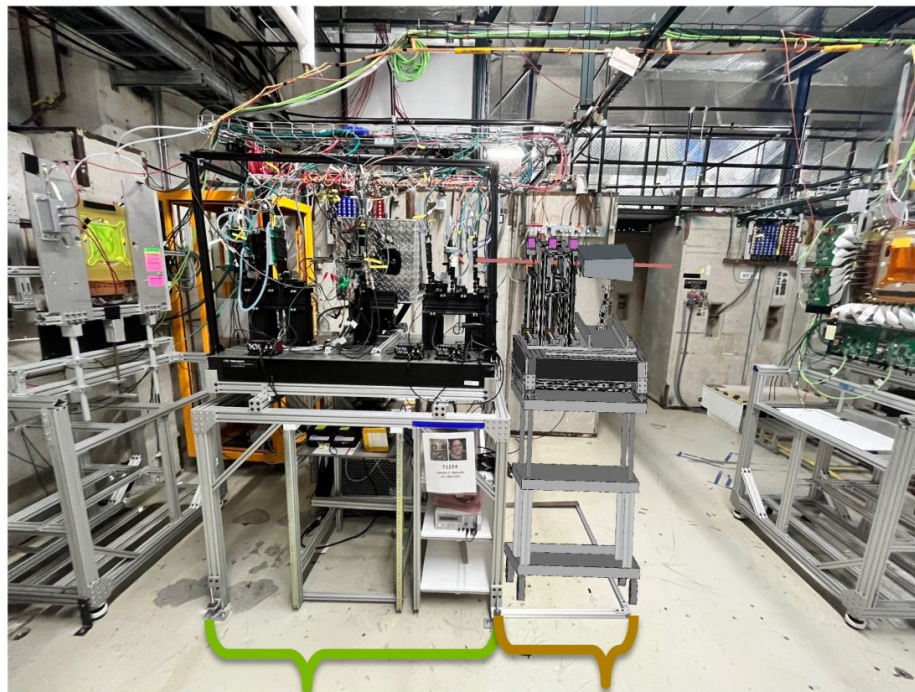
- Physics benchmark of energy response to pions
- Physics benchmark of  $e/\pi$  separation
- Technical benchmark of streaming readout of both technologies

**How performance of modern family of SiPMs improves the SciFi/Pb part response wrt the GlueX BCAL response?**

- Benchmark light response and calibrate simulations
- Impact on future design studies related to usage of optical cookies, shape of lightguides, etc.
  - **Photon Detection Efficiency** for GlueX SiPMs (Hamamatsu S12045(X)): **~33%**
  - Modern family of SiPMs (e.g. s14160/14161): **~50%** (see backup slides 18-20)

# GENERAL R&D SETUP AT FNAL

- Add BIC prototype calorimeter behind existing Argonne ATLAS Pixel telescope with AstroPix setup at MTest
- Rotating stage to simulate particles incident at angles up to  $45^\circ$  ( $\eta \sim 1$ )
- Ability to lower BIC setup out of the beam, no need to uninstall for other experiments to run
  - Proximity to Argonne enables occasional opportunistic running



Current ANL AstroPix  
Telescope Setup

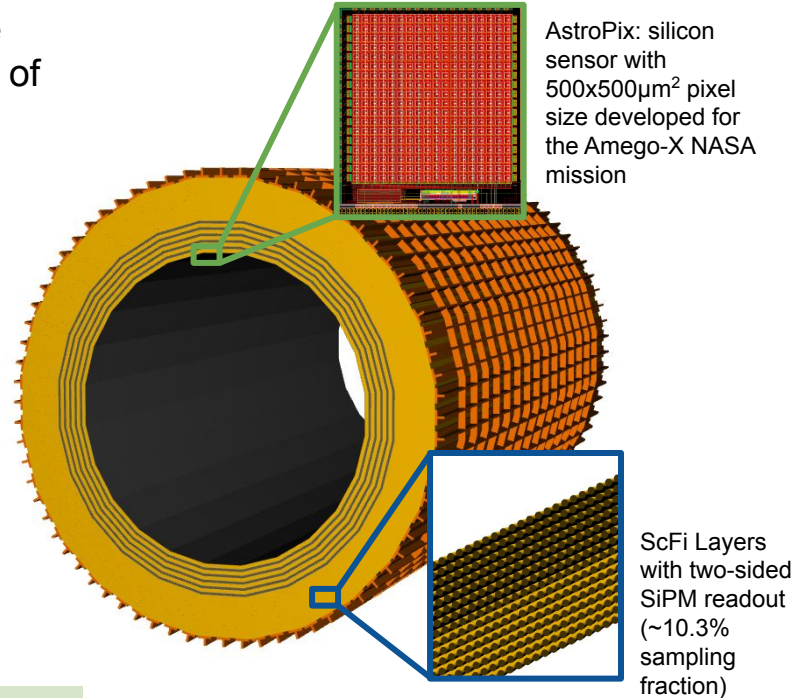
Planned BIC Setup

# SUMMARY

## Addressing the unique challenges for the barrel region in ePIC

**Hybrid concept:** 4 (+2) layers of Astropix interleaved with the first 5 Pb/ScFi layers, followed by a large volume with the rest of the Pb/ScFi layers

- ✓ Deep calorimeter ( $\eta = 0 \sim 17.1 X_0$ ) while compact at  $\sim 40$  cm
- ✓ Excellent energy resolution ( $5.2\% / \sqrt{E} \oplus 1.0\%$ )
- ✓ Unrivaled low-energy electron-pion separation by combining the energy measurement with shower imaging
- ✓ Unrivaled position resolution due to the silicon layers
- ✓ Deep enough to serve as inner HCal
- ✓ Very good low-energy performance
- ✓ Wealth of information enables new measurements, ideally suited for particle-flow
- ✓ Serves as tracking layer behind the DIRC



Checks all the boxes!