The $\mu \rightarrow e\gamma$ decay in an EW-scale non-sterile right-handed neutrino model

Dinh Nguyen Dinh

Institute of Physics, Hanoi, Vietnam

Online workshop: Darkness on the table, 2021



Introduction ●○○	An EW-scale ν_R model 00000000	The $\mu \rightarrow e\gamma$ decay 000000000	Conclusi 00

- Neutrino oscillation discovery is evidence of massive and missing neutrinos; also lepton flavor violation (LFV). However, it has not been determined in the charged lepton sector.
- In minimal extended models, LFV process branching ratios in charged lepton sector is extremely tiny; $Br(\mu \rightarrow e\gamma) \sim 10^{-55}$.
- Current and future expected sensitivities by MEG:

$$\begin{array}{lll} {\rm BR}(\mu^+ \to e^+ \gamma) &< & 4.2 \times 10^{-13} \,, \\ {\rm BR}(\mu^+ \to e^+ \gamma) &< & 6.0 \times 10^{-14}. \end{array}$$

 \implies Not possible to determine in practice!



MEG: Eur. Phys. J., Vol. C76, No. 8, P.434, 2016; Arxiv:2107.10767



The $\mu \rightarrow e\gamma$ decay 00000000

Introduction

- There are more experimental confirmations prove that SM requires to be modified beside massive neutrino problem, such as dark matter, matter anti-matter asymmetry, ...
- In those extensions, the LFV branching ratios might be much higher and reach the experimental sensitivities.
- In this report, we discuss $\mu \rightarrow e\gamma$ decay in a EW-scale ν_R model
 - Derive the branching ratio
 - Set constraints on relevant parameters, predict observable possibilities by current and future experiments.



Outline

- Briefly introduce the model
- ② Calculate the form factors and derive decay branching ratio
- Discussions and set constraints on involving couplings, for the cases of one-loop diagrams involving
 - W boson
 - singly charged scalars
 - Heavy neutral scalars
 - Light neutral scalar
- Conclusion





Introduction	An EW-scale ν_R model	The $\mu \rightarrow e\gamma$ decay 00000000	Conclusion
000	••••••••		00

Model overview

- The model is constructed based on a symmetric group $SU(2) \times U(1)_Y \times U(1)_{SM} \times U(1)_{MF}$, where $SU(2) \times U(1)_Y$ is gauge group, and $U(1)_{SM} \times U(1)_{MF}$ is a global symmetry.
- The particle contents are

SM particlesMirror partners
$$\ell_L = (\nu_L, e_L)^T, e_R$$
 $\ell_R^M = (\nu_R, e_R^M)^T, e_L^M$ $q_L = (u_L, d_L)^T, u_R, d_R$ $q_R^M = (u_R^M, d_R^M)^T, u_L^M, d_L^M$

• The Higgs sector consists of:

- Two Higgs doublets: $\Phi_2 = (\phi_2^+, \phi_2^0), \Phi_{2M} = (\phi_{2M}^+, \phi_{2M}^0),$
- A Y = 0 real triplet $\xi = (\xi^+, \overline{\xi}^0, \xi^-)$, and a Y = 2 complex triplet $\tilde{\chi}$

$$\tilde{\chi} = \frac{1}{\sqrt{2}} \vec{\tau} \cdot \vec{\chi} = \begin{pmatrix} \frac{1}{\sqrt{2}} \chi^+ & \chi^{++} \\ \chi^0 & -\frac{1}{\sqrt{2}} \chi^+ \end{pmatrix}$$



• A Higgs singlet ϕ_S .

An EW-scale ν_R model	The $\mu ightarrow e \gamma$ decay	Conclusion
0000000		

Model overview

• Global symmetry

- Global symmetry $U(1)_{SM} \times U(1)_{MF}$ is imposed to prevent some unexpected couplings.
- Transformations of matter and Higgs fields are defined as following:
 i) U(1)_{SM}: Ψ = {Φ₂, qSM_L, ℓSM_L} transforms as Ψ → e^{iα_{SM}}Ψ,
 ii) U(1)_{MF}: Ψ = {Φ_{2M}, q^M_R, ℓ^M_R} transforms as Ψ → e^{iα_{MF}}Ψ,
 iii) φ_S → e^{-i(α_{MF}-α_{SM})}φ_S, χ̃ → e^{-2iα_{MF}} χ̃,
- The unmentioned fields are singlets.
- The Yukawa couplings are:

$$\mathcal{L}_{Y}^{\ell} = g_{\ell} \bar{\ell}_{L} \Phi_{2} e_{R} + g_{\ell}^{M} \bar{\ell}_{R}^{M} \Phi_{2M} e_{L}^{M} + g_{\ell s} \bar{\ell}_{L} \phi_{s} \ell_{R}^{M} + h.c.$$

 $\mathcal{L}_Y^q = g_u \bar{q}_L \tilde{\Phi}_2 u_R + g_d \bar{q}_L \Phi_2 d_R + g_u^M \bar{q}_R^M \tilde{\Phi}_{2M} u_L^M + g_d^M \bar{q}_R^M \Phi_{2M} d_L^M + g_{qs} \bar{q}_L \phi_S q_R^M + h.c.$

$$\mathcal{L}_{\nu_R} = g_M l_R^{M,T} \, \sigma_2 \, \tilde{\chi} \, l_R^M \,,$$

where σ_2 is the second Pauli matrix, $\tilde{\Phi}_2 = i\sigma_2 \Phi_2^*$ and $\tilde{\Phi}_{2M} = i\sigma_2 \Phi_{2M}^*$.



An EW-scale ν_R model 0000000

The $\mu \rightarrow e\gamma$ decay 00000000

Conclusion 00

Symmetry breaking

- Higgs fiels develop their VEVs as: $\langle \Phi_2 \rangle = (0, v_2/\sqrt{2})^T$, $\langle \Phi_{2M} \rangle = (0, v_{2M}/\sqrt{2})^T$, $\langle \chi^0 \rangle = v_M$, and $\langle \phi_S \rangle = v_S$,
- Charged lepton masses:
 - The mass matrix

$$M_\ell = \left(egin{array}{cc} m_\ell & m_\ell^D \ (m_\ell^D)^\dagger & m_{\ell M} \end{array}
ight) \, ,$$

where $m_{\nu}^{D} = m_{\ell}^{D} = g_{\ell s} v_{s}$, $m_{\ell} = g_{\ell} v_{2} / \sqrt{2}$, and $m_{\ell M} = g_{\ell}^{M} v_{2M} / \sqrt{2}$. • For $m_{\ell}^{D} \ll m_{\ell M}$, $m_{\ell}^{D} \ll m_{\ell}$

$$M_\ell = \left(egin{array}{cc} m_\ell & m_\ell^D \ (m_\ell^D)^\dagger & m_{\ell M} \end{array}
ight) = \left(egin{array}{cc} I & R_\ell \ -R_\ell^\dagger & I \end{array}
ight)^\dagger \left(egin{array}{cc} ilde{m}_\ell & 0 \ 0 & ilde{m}_{\ell M} \end{array}
ight) \left(egin{array}{cc} I & R_\ell \ -R_\ell^\dagger & I \end{array}
ight),$$

where $R_{\ell} \approx \frac{m_{\ell}^{D}}{m_{\ell M}} \ll 1$, $\tilde{m}_{\ell} \approx m_{\ell}$, $\tilde{m}_{\ell M} \approx m_{\ell M}$.

• If we express $\tilde{m}_{\ell} = U_{\ell L} m_{\ell}^{d} U_{\ell R}^{\dagger}, \tilde{m}_{\ell M} = U_{\ell L}^{M} m_{\ell M}^{d} U_{\ell R}^{M^{\dagger}}$, then

$$\begin{pmatrix} \ell_{L(R)} \\ \ell_{L(R)}^{M} \end{pmatrix} = \begin{pmatrix} U_{\ell L(R)} & -R_{\ell} U_{\ell L(R)}^{M} \\ R_{\ell}^{\dagger} U_{\ell L(R)} & U_{\ell L(R)}^{M} \end{pmatrix} \begin{pmatrix} \ell_{L(R)}' \\ \ell_{L(R)}^{M} \end{pmatrix} .$$



The $\mu \rightarrow e\gamma$ decay 00000000

Symmetry breaking

- Neutrino masses:
 - The mass matrix

$$M_{\nu} = \begin{pmatrix} 0 & m_{\nu}^D \\ (m_{\nu}^D)^T & M_R \end{pmatrix} . \tag{1}$$

where $m_{\nu}^D = m_{\ell}^D = g_{\ell s} v_S$, $M_R = g_M v_M$.

• For $m_{\nu}^D \ll M_R$, we can approximately blocked diagonalized

$$M_{\nu} = \begin{pmatrix} 0 & m_{\nu}^{D} \\ (m_{\nu}^{D})^{T} & M_{R} \end{pmatrix} = \begin{pmatrix} I & R_{\nu} \\ -R_{\nu}^{\dagger} & I \end{pmatrix}^{T} \begin{pmatrix} \tilde{m}_{\nu} & 0 \\ 0 & \tilde{m}_{\nu R} \end{pmatrix} \begin{pmatrix} I & R_{\nu} \\ -R_{\nu}^{\dagger} & I \end{pmatrix},$$

where $R_{\nu} \approx \frac{m_{\nu}^{D}}{M_{R}}$, and

 $\left(\begin{array}{c}\nu_L\\(\nu_R)^c\end{array}\right) = \left(\begin{array}{c}U_\nu & -R_\nu U_\nu^M\\R_\nu^\dagger U_\nu & U_\nu^M\end{array}\right) \left(\begin{array}{c}\chi_\nu\\\mathcal{X}_M\end{array}\right),$

$$\tilde{m}_{\nu} \approx -\frac{(m_{\nu}^D)^2}{M_R} = -\frac{(g_{\ell s} v_S)^2}{g_M v_M}, \quad \tilde{m}_{\nu R} \approx M_R.$$
(2)

• Suppose that $\tilde{m}_{\nu} = U_{\nu}^* m_{\nu}^d U_{\nu}^{\dagger}, \tilde{m}_{\nu R} = U_{\nu}^{M^*} m_{\nu M}^d U_{\nu}^{M^{\dagger}}$, then



D. N. Dinh (IOP-VAST)

The $\mu \rightarrow e\gamma$ decay 00000000

Symmetry breaking in Higgs sector

- If there is only one Higgs triplet, the quantity $\rho = M_w/(M_z \cos \theta_w) = 1$ might be spoiled out, thus Higgs triplets ξ , $\tilde{\chi}$ are introduced.
- It is proved in M.S. Chanowitz et. al., Phys. Lett. B 165 (1985) 105 that if the two triplets combined to form (3, 3) representation under global $SU(2)_L \times SU(2)_R$ of the Higgs potential; After gaining VEV, the symmetry is broken downto a custodial SU(2), then $\rho = 1$.
- In this model, two triplets and doublets form (3,3) and (2,2) representations as:

$$\chi = \begin{pmatrix} \chi^0 & \xi^+ & \chi^{++} \\ \chi^- & \xi^0 & \chi^+ \\ \chi^{--} & \xi^- & \chi^{0*} \end{pmatrix},$$
$$\Phi_2 = \begin{pmatrix} \phi_{2}^{0,*} & \phi_{2}^+ \\ \phi_{2}^- & \phi_{2}^0 \end{pmatrix}, \quad \Phi_{2M} = \begin{pmatrix} \phi_{2M}^{0,*} & \phi_{2M}^+ \\ \phi_{2M}^- & \phi_{2M}^0 \end{pmatrix}.$$

• $\langle \Phi_2^0 \rangle = (v_2/\sqrt{2}), \langle \Phi_{2M}^0 \rangle = (v_{2M}/\sqrt{2}) \text{ and } \langle \chi^0 \rangle = v_M,$



The $\mu \rightarrow e\gamma$ decay 00000000

Symmetry breaking in Higgs sector

• Generated physical scalars as representations of the global custodial symmetry as:

five-plet (quintet) $\rightarrow H_5^{\pm\pm}, H_5^{\pm}, H_5^{0};$ triplet $\rightarrow H_3^{\pm}, H_3^{0};$ triplet $\rightarrow H_{3M}^{\pm}, H_{3M}^{0};$ three singlets $\rightarrow H_1^{0}, H_{1M}^{0}, H_1^{0'}.$

- Three singlets H_1^0 , H_{1M}^0 , $H_1^{0\prime}$ are not mass eigenstates; $H_1^0 = \sum_i^3 \alpha_i \tilde{H}_i$, $H_{1M}^0 = \sum_i^3 \alpha_i^M \tilde{H}_i$, where $\sum_i^3 |\alpha_i|^2 = 1$ and $\sum_i^3 |\alpha_i^M|^2 = 1$.
- The light singlet ϕ_s^0 .



An EW-scale ν_R model 0000000

The $\mu \rightarrow e\gamma$ decay 00000000

LFV vertexes

	Vertices	Couplings
	$(\bar{e}'_L \gamma^\mu \chi_L) W^\mu$	$-i \frac{g}{\sqrt{2}} U^L_{W\mu} = -i \frac{g}{\sqrt{2}} U_{PMNS}$
	$(\bar{e}'_L \gamma^\mu \chi^M_L) W^\mu$	$i\frac{g}{\sqrt{2}}U^{ML}_{W\mu} = i\frac{g}{\sqrt{2}}\tilde{R}_{\nu}\left(U^{M}_{PMNS}\right)^{*}$
	$(\bar{e}_R^{\prime}\gamma^{\mu}\chi_L^c)W_{\mu}^{-}$	$-i\frac{g}{\sqrt{2}}U^{\vec{R}}_{W\mu} = -i\frac{g}{\sqrt{2}}\tilde{R}^{T}_{\nu}(U_{PMNS})^{*}$
	$\bar{e}'_R \chi_L H_3^-$	$-i\frac{g}{2}Y^{L}_{H_{3}} = -i\frac{g^{s}M}{2M_{W}c_{M}}m^{d}_{\ell}U_{PMNS}$
	$\bar{e}_R' \chi_L^M H_3^-$	$i\frac{g}{2}Y_{H_{3}}^{ML} = i\frac{gs_{M}}{2M_{W}c_{M}}m_{\ell}^{d}\tilde{R}_{\nu}\left(U_{PMNS}^{M}\right)^{*}$
٩	$\bar{e}_L' \chi_L^{Mc} H_3^-$	$-i\frac{g}{2}\tilde{Y}_{H_{3}}^{MR} = -i\frac{g}{2M_{W}c_{M}}\tilde{R}_{\ell}m_{\ell M}^{d}U_{PMNS}^{M}$
	$\bar{e}'_R \chi_L H_{3M}^-$	$-i\frac{g}{2}Y_{H_{3M}}^{L} = -i\frac{g}{2M}\frac{s}{2M}m_{\ell}^{d}U_{PMNS}$
	$\bar{e}_R' \chi_L^M H_{3M}^-$	$i\frac{g}{2}Y_{H_{3M}}^{ML} = i\frac{g}{2M_Ws_2c_M}m_\ell^d\tilde{R}_\nu \left(U_{PMNS}^M\right)^*$
	$\bar{e}_L' \chi_L^{Mc} H_{3M}^-$	$-i\frac{g}{2}Y_{H_{3M}}^{\widetilde{MR}} = -i\frac{g}{2M_Ws_{2M}c_M}\tilde{R}_\ell m_{\ell M}^d U_{PMNS}^M$
	$\bar{e}_{R}^{\prime}e_{L}^{M\prime}\phi_{s}^{0}$	$-i\frac{g}{2}Y^{ML}_{\phi^0_S} = -iU^{\dagger}_{\ell R}g_{\ell s}U^M_{\ell L}$
	$\overline{e}_L' e_R^{M\prime} \phi_s^0$	$-i\frac{g}{2}Y_{\phi_{c}}^{MR} = -iU_{\ell L}^{\dagger}g_{\ell s}U_{\ell R}^{M}$

Tabelle: Vertices that contribute to the decay rates $\ell \to \ell' \gamma$.



IAFOSTEL

Introd	

An EW-scale ν_R model

The $\mu \rightarrow e\gamma$ decay 00000000

LFV vertexes

۲

$$\begin{split} & \left(\tilde{e}'_{R} e^{M'}_{L} \tilde{H}^{0}_{i} \right) & -i \frac{g}{2} Y^{ML}_{\tilde{H}^{0}_{i}} = -i \frac{g}{2M_{W}} \left[\frac{\alpha_{i}}{s_{2}} m^{d}_{\ell} \tilde{R}_{\ell} + \frac{\alpha^{M}_{i}}{s_{2M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right], \\ & \left(\tilde{e}'_{L} e^{M'}_{R} \tilde{H}^{0}_{i} \right) & -i \frac{g}{2} Y^{MR}_{\tilde{H}^{0}_{i}} = -i \frac{g}{2M_{W}} \left[\frac{\alpha_{i}}{s_{2}} m^{d}_{\ell} \tilde{R}_{\ell} + \frac{\alpha^{M}_{i}}{s_{2M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right], \\ & \left(\tilde{e}'_{L} e^{M'}_{R} H^{0}_{3} \right) & -i \frac{g}{2} Y^{ML}_{H^{0}_{3}} = -i \frac{g}{2M_{W}} \left[\frac{s_{M}}{s_{M}} m^{d}_{\ell} \tilde{R}_{\ell} + \frac{s_{M}}{c_{M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right], \\ & \left(\tilde{e}'_{L} e^{M'}_{R} H^{0}_{3} \right) & -i \frac{g}{2} Y^{MR}_{H^{0}_{3}} = -i \frac{g}{2M_{W}} \left[-\frac{s_{M}}{c_{M}} m^{d}_{\ell} \tilde{R}_{\ell} - \frac{s_{M}}{c_{M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right], \\ & \left(\tilde{e}'_{L} e^{M'}_{R} H^{0}_{3M} \right) & -i \frac{g}{2} Y^{ML}_{H^{0}_{3M}} = -i \frac{g}{2M_{W}} \left[-\frac{s_{2M}}{s_{2}} m^{d}_{\ell} \tilde{R}_{\ell} - \frac{s_{2}}{s_{2M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right], \\ & \left(\tilde{e}'_{L} e^{M'}_{R} H^{0}_{3M} \right) & -i \frac{g}{2} Y^{ML}_{H^{0}_{3M}} = -i \frac{g}{2M_{W}} \left[-\frac{s_{2M}}{s_{2}} m^{d}_{\ell} \tilde{R}_{\ell} + \frac{s_{2}}{s_{2M}} \tilde{R}_{\ell} m^{d}_{\ell M} \right]. \end{split}$$

Here the notations $U_{PMNS} = U_{\ell}^{\dagger} U_{\nu}$, which is the famous neutrino mixing PMNS matrix, $U_{PMNS}^{M} = U_{\ell}^{M\dagger} U_{\nu}^{M}$ and $\tilde{R}_{\ell(\nu)} = U_{\ell}^{\dagger} R_{\ell(\nu)} U_{\ell}^{M}$ have been used. For simplicity, we have also neglected the complex phases in U_{ℓ} and U_{ℓ}^{M} .

An EW-scale ν_R mo 00000000 The $\mu \rightarrow e\gamma$ decay •••••••

Formfactors & $Br(\mu \rightarrow e + \gamma)$

• The effective Lagrangian has form:

$$\mathcal{L}_{eff} = -4 \frac{eG_F}{\sqrt{2}} m_\mu \left(A_R \bar{e} \sigma_{\mu\nu} P_R \mu + A_L \bar{e} \sigma_{\mu\nu} P_L \mu \right) F^{\mu\nu} + H.c.$$

$$\begin{split} A_{R} &= -\sum_{H^{Q},k} \frac{M_{W}^{2}}{64\pi^{2}M_{H}^{2}} \left[\left(Y_{H}^{L} \right)_{\mu k} \left(Y_{H}^{L} \right)_{ek}^{*} G_{H}^{Q}(\lambda_{k}) + \frac{m_{k}}{m_{\mu}} \left(Y_{H}^{R} \right)_{\mu k} \left(Y_{H}^{L} \right)_{ek}^{*} \times R_{H}^{Q}(\lambda_{k}) \right] \\ &+ \frac{1}{32\pi^{2}} \sum_{k} \left[\left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) - \left(U_{W\mu}^{R} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} R_{\gamma}(\lambda_{k}) \right], \end{split}$$

$$\begin{split} A_{L} &= -\sum_{H^{Q},k} \frac{M_{W}^{2}}{64\pi^{2}M_{H}^{2}} \left[\left(Y_{H}^{R} \right)_{\mu k} \left(Y_{H}^{R} \right)_{ek}^{*} G_{H}^{Q}(\lambda_{k}) + \frac{m_{k}}{m_{\mu}} \left(Y_{H}^{L} \right)_{\mu k} \left(Y_{H}^{R} \right)_{ek}^{*} R_{H}^{Q}(\lambda_{k}) \right] \\ &+ \frac{1}{32\pi^{2}} \sum_{k} \left[\left(U_{W\mu}^{R} \right)_{\mu k} \left(U_{W\mu}^{R} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) - \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{R} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} R_{\gamma}(\lambda_{k}) \right], \end{split}$$

where $H^Q = \phi_S^0, \tilde{H}_i^0$ $(i = 1, 2, 3), H_3^0, H_{3M}^0, H_3^+, H_{3M}^+$, and m_k are the masses of associated fermions that along with either H^Q or W_μ to form loops, and $\lambda_k = m_k^2/M_{W_\mu(H^Q)}^2$.

D. N. Dinh (IOP-VAST)

An EW-scale ν_R mo

The $\mu \rightarrow e\gamma$ decay $0 \bullet 0000000$

Formfactors & $Br(\mu \rightarrow e + \gamma)$

• The branching ratio of $\mu \rightarrow e + \gamma$ decay is easily obtained as

$$\operatorname{Br}(\mu \to e + \gamma) = 384\pi^2 (4\pi\alpha_{em}) \left(|A_R|^2 + |A_L|^2 \right),$$

where $\alpha_{em} = 1/137$ is the fine-structure constant.

• Functions $G_H^Q(x)$, $R_H^Q(x)$, $G_{\gamma}(x)$, and $R_{\gamma}(x)$ are defined as:

$$G_{H}^{Q}(x) = -\frac{(3Q-1)x^{2}+5x-3Q+2}{12(x-1)^{3}} + \frac{1}{2}\frac{x(Qx-Q+1)}{2(x-1)^{4}}\log(x),$$
(3)

$$R_{H}^{Q}(x) = \frac{(2Q-1)x^{2} - 4(Q-1)x + 2Q - 3}{2(x-1)^{3}} - \frac{Qx - (Q-1)}{(x-1)^{3}}\log(x),$$
(4)

$$G_{\gamma}(x) = \frac{20x^2 - 7x + 2}{4(x-1)^3} - \frac{3}{2} \frac{x^3}{(x-1)^4} \log(x),$$
(5)

$$R_{\gamma}(x) = -\frac{x^2 + x - 8}{2(x - 1)^2} + \frac{3x(x - 2)}{(x - 1)^3}\log(x).$$
(6)



The $\mu \rightarrow e\gamma$ decay 0000000

$\nu - w$ 1-loop diagrams

• Contribution of the light neutrinos to the process are extremely small

$$\begin{split} \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) \approx \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(0) = \left(U_{W\mu}^{L\dagger} U_{W\mu}^{L} \right)_{e\mu} G_{\gamma}(0) \approx 0. \\ \\ \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{R} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} \sim 10^{-9} \end{split}$$





The $\mu \rightarrow e\gamma$ decay 0000000

$\nu - w$ 1-loop diagrams

• Contribution of the light neutrinos to the process are extremely small

$$\begin{split} \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) \approx \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(0) = \left(U_{W\mu}^{L\dagger} U_{W\mu}^{L} \right)_{e\mu} G_{\gamma}(0) \approx 0. \\ \\ \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{R} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} \sim 10^{-9} \end{split}$$

 $\nu - w \text{ loop contribution}$ $\left(U_{W_{\mu}}^{L} \right)_{\mu k} \left(U_{W_{\mu}}^{L} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) - \left(U_{W_{\mu}}^{L} \right)_{\mu k} \left(U_{W_{\mu}}^{R} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} R_{\gamma}(\lambda_{k})$





The $\mu \rightarrow e\gamma$ decay 00000000

$\nu - w$ 1-loop diagrams

• Contribution of the light neutrinos to the process are extremely small

$$\sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(\lambda_{k}) \approx \sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{L} \right)_{ek}^{*} G_{\gamma}(0) = \left(U_{W\mu}^{L\dagger} U_{W\mu}^{L} \right)_{e\mu} G_{\gamma}(0) \approx 0.$$
$$\sum_{k=1}^{3} \left(U_{W\mu}^{L} \right)_{\mu k} \left(U_{W\mu}^{R} \right)_{ek}^{*} \frac{m_{k}}{m_{\mu}} \sim 10^{-9}$$

• Heavy neutrinos



The $\mu \rightarrow e\gamma$ decay 00000000

$\nu - H^-$ loop diagrams

• The interfered terms give dominated contributions

 $\frac{m_k}{m_\mu} \sim \frac{100 \text{ GeV}}{100 \text{ MeV}} \sim 10^3.$





The $\mu \rightarrow e\gamma$ decay 00000000

$\nu - H^-$ loop diagrams

• The interfered terms give dominated contributions

$$\frac{m_k}{m_\mu} \sim \frac{100 \text{ GeV}}{100 \text{ MeV}} \sim 10^3.$$

 $\nu - H^{-} \text{ loop contribution}$ $\left(Y_{H}^{L}\right)_{\mu k} \left(Y_{H}^{L}\right)_{e k}^{*} G_{H}^{Q}(\lambda_{k}) + \frac{m_{k}}{m_{\mu}} \left(Y_{H}^{R}\right)_{\mu k} \left(Y_{H}^{L}\right)_{e k}^{*} \times R_{H}^{Q}(\lambda_{k})$



The $\mu \rightarrow e\gamma$ decay 00000000

$\nu - H^-$ loop diagrams

• The interfered terms give dominated contributions

 $\frac{m_k}{m_\mu} \sim \frac{100 \text{ GeV}}{100 \text{ MeV}} \sim 10^3.$





The $\mu \rightarrow e\gamma$ decay 00000000

$\nu - H^-$ loop diagrams

- The interfered terms give dominated contributions
- From the figure, we have:

 $|(Y_{H^{-}}^{ML})^{\dagger}Y_{H^{-}}^{MR}|_{\mu e}^{2} + |(Y_{H^{-}}^{MR})^{\dagger}Y_{H^{-}}^{ML}|_{\mu e}^{2} \lesssim 3.0 \times 10^{-16} (6.0 \times 10^{-16}) \text{ for } m_{\chi_{L}^{M}} = 80 \ (200) \text{ GeV},$

 $|(Y_{H^-}^{ML})^{\dagger}Y_{H^-}^{MR}|_{\mu e}^2 + |(Y_{H^-}^{MR})^{\dagger}Y_{H^-}^{ML}|_{\mu e}^2 \lesssim 4.0 \times 10^{-17} (8.0 \times 10^{-17}) \text{ for } m_{\chi_L^M} = 80 \ (200) \text{ GeV}.$



$$H_{3}^{-}: \left| Y_{H_{3}}^{L\dagger} Y_{H_{3M}}^{R} \right|_{21}^{2} + \left| Y_{H_{3M}}^{R\dagger} Y_{H_{3}}^{L} \right|_{21}^{2} \sim 2.2 \times 10^{-29} \left(\frac{s_{M}^{2}}{c_{M}^{4}} \right)^{2}$$
$$H_{3M}^{-}: \left| Y_{H_{3M}}^{L\dagger} Y_{H_{3M}}^{R} \right|_{21}^{2} + \left| Y_{H_{3M}}^{R\dagger} Y_{H_{3M}}^{L} \right|_{21}^{2} \sim 2.2 \times 10^{-29} \left(\frac{s_{M}}{s_{2}c_{M}^{2}} \right)^{2}$$



An EW-scale ν_R mod 00000000 The $\mu \rightarrow e\gamma$ decay 00000000

$\ell^M - H^0$ loop diagrams



• Constraints obtained from the left panel

$$\begin{split} & 2|(Y_{H0}^{ML})^{\dagger}Y_{H0}^{MR}|^{2}_{\mu e} \lesssim 7.0 \times 10^{-17} (3.0 \times 10^{-16}) \text{ for } m_{\ell}^{M} = 80 \ (200) \text{ GeV (green lines)} \\ & 2|(Y_{H0}^{ML})^{\dagger}Y_{H0}^{MR}|^{2}_{\mu e} \lesssim 1.0 \times 10^{-17} (4.5 \times 10^{-17}) \text{ for } m_{\ell}^{M} = 80 \ (200) \text{ GeV (red lines)} \end{split}$$

• Theoretical estimations

$$2\left|Y_{H^{0}}^{L^{\dagger}}Y_{H^{0}}^{R}\right|_{21}^{2} = 2\alpha^{4} \times \frac{\left|\tilde{R}_{\ell}(m_{\ell M}^{d})^{2}\tilde{R}_{\ell}\right|_{21}^{2}}{M_{W}^{4}} \sim 4.4 \times 10^{-23} \alpha^{4},$$

IOP

where
$$\alpha$$
 stands for $\frac{\alpha_i}{s_2}$, $\frac{s_M}{c_M}$ or $\frac{s_2}{s_{2M}}$, corresponding to \tilde{H}_i^0 , $(i = 1, 2, 3)$, H_0^0 or H_0^0 , respectively interval \tilde{H}_i^0 , $(i = 1, 2, 3)$, H_0^0 or H_0^0 , respectively \tilde{H}_i^0 , \tilde{H}_i^0 ,

An EW-scale ν_R mo

The $\mu \rightarrow e\gamma$ decay 00000000

$\ell^M - H^0$ loop diagrams

• Constraints obtained from the left panel

$$2|(Y_{H0}^{ML})^{\dagger}Y_{H0}^{MR}|_{\mu e}^{2} \lesssim 7.0 \times 10^{-17} (3.0 \times 10^{-16}) \text{ for } m_{\ell}^{M} = 80 (200) \text{ GeV (green lines)}$$
$$2|(Y_{H0}^{ML})^{\dagger}Y_{H0}^{MR}|_{\mu e}^{2} \lesssim 1.0 \times 10^{-17} (4.5 \times 10^{-17}) \text{ for } m_{\ell}^{M} = 80 (200) \text{ GeV (red lines)}$$

Theoretical estimations

$$2 \left| Y_{H^0}^{L\dagger} Y_{H^0}^R \right|_{21}^2 = 2\alpha^4 \times \frac{ \left| \tilde{R}_\ell(m_{\ell M}^d)^2 \tilde{R}_\ell \right|_{21}^2}{M_W^4} \sim 4.4 \times 10^{-23} \alpha^4,$$

where α stands for $\frac{\alpha_i}{s_2}$, $\frac{s_M}{c_M}$ or $\frac{s_2}{s_{2M}}$, corresponding to \tilde{H}_i^0 , (i = 1, 2, 3), H_3^0 or H_{3M}^0 , respectively. In the H_3^0 case, for $c_M = 0.01$, the factor $2 \left| Y_{H^0}^{L^{\dagger}} Y_{H^0}^R \right|_{21}^2 \sim 4.4 \times 10^{-15}$, which is about two order higher than the expected sensitive limit of the future experiment. In fact, our calculation show that the factor might be lager than 10^{-17} for $c_M \leq 0.03$.



An EW-scale ν_R mo

The $\mu \rightarrow e\gamma$ decay 0000000

$\ell^M - \phi^0_s$ loop diagrams



• One obtains from the left panel

$$\begin{split} & 2|(Y^L_{\phi^{+}_{\mathcal{S}}})^{\dagger}Y^R_{\phi^{+}_{\mathcal{S}}}|^2_{\mu e} \lesssim 4.7 \times 10^{-17} (3.0 \times 10^{-16}) \text{ for } m^M_{\ell} = 80 \ (200) \text{ GeV}, \\ & 2|(Y^L_{\phi^{+}_{\mathcal{S}}})^{\dagger}Y^R_{\phi^{+}_{\mathcal{S}}}|^2_{\mu e} \lesssim 6.5 \times 10^{-18} (4.2 \times 10^{-17}) \text{ for } m^M_{\ell} = 80 \ (200) \text{ GeV}. \end{split}$$

• Theoretical estimations

 $\begin{aligned} |g_{\ell s}| \lesssim 2.3 \times 10^{-5} (3.6 \times 10^{-5}) \text{ for } m_{\ell}^{M} = 80 \ (200) \text{ GeV} \\ |g_{\ell s}| \lesssim 1.4 \times 10^{-5} (2.3 \times 10^{-5}) \text{ for } m_{\ell}^{M} = 80 \ (200) \text{ GeV} \end{aligned} \right\} \implies v_{s} \text{ at GeV order or higher.}$



D. N. Dinh (IOP-VAST)

Conclusion

- We have introduce the model, calculate the form-factor and derive the $\mu \rightarrow e\gamma$ decay branching ratio.
- Contribution of the νW_{μ} loop diagrams is less important and does not give any meaning constraint on the model's parameters.
- The branching ratio might reach the future experimental sensitivities due to the main contributions from heavy Higgs scalar loop diagrams, in which the neutral scalar channels are the most important.
- The couplings $g_{\ell s}$ is constrained to be in order of 10^{-5} , which implies that v_s magnitude is in range of few GeVs to several ten GeVs.



Thanks you!

