Vector meson mass in the chiral symmetry restored vacuum

Su Houng Lee



- 1. General remarks on hadron mass
- 2. Vector meson in the chiral symmetry restored vacuum
- 3. K_1 and K^* in nuclear matter

4. Summary

Previous work +

- T. Song, T. Hatsuda, Su Houng Lee, PLB792 (2019) 160
- Jisu Kim and Su Houng Lee, PRD103 (2021) L051501 + in preparation
- Haesom Sung, et al. PLB819(2021)136388

Confinement and hadron mass (Hashimoto, Miyamura, Hirose, Kanki)

VOLUME 57, NUMBER 17

27 OCTOBER 1986

Mass Shift of Charmonium near Deconfining Temperature and Possible Detection in Lepton-Pair Production





Chiral symmetry breaking and hadron mass

PHYSICAL REVIEW D 93, 054035 (2016)

Mass of heavy-light mesons in a constituent quark picture with partially restored chiral symmetry

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Gauge invariant + relativistic method:

Methods based on correlation functions

- definitions and effects on masses

- 1. Confinement
- 2. Chiral symmetry breaking
- 3. $U_A(1)$ effect

Confinement and gluon condensates

Wilson Loops and potential



Local operators

OPE for Wilson lines: Shifman NPB73 (80) Dosch, Simonov PLB339 (88)

W(S-T) = 1- $\langle \alpha / \pi E^2 \rangle$ (ST)² +...

$$W(S-S) = 1- \langle \alpha / \pi B^2 \rangle (SS)^2 + \dots$$

SHLee PRD40 (89): ------Non-perturbative Gluon condensate above Tc

Manousakis, Polonyi PRL 1987



Morita, SHLee PRL 2008, PRD 2009



Chiral symmetry breaking (m→0) : order parameter

• Quark condensate $SU(N_F)_L \times SU(N_F)_R \to SU(N_F)_V$ $\langle \overline{q}q \rangle = \langle \overline{q}_L q_R + \overline{q}_R q_L \rangle = -\lim_{x \to 0} \langle \operatorname{Tr}[S(x,0)] \rangle = -\lim_{x \to 0} \langle \frac{1}{2} \operatorname{Tr}[S(x,0) - i\gamma^5 S(x,0) i\gamma^5] \rangle$ $q_L \to q_R$ Chiral rotation $q \to \exp(i\gamma^5 \tau^a \alpha^a) q$

Casher Banks formula: nontrivial zero mode ($\lambda = 0$) contribution

 $\left\langle \frac{\alpha}{\pi} G^2 \right\rangle = 12 \left\langle \sum_{\lambda} \rho(\lambda) \right\rangle$

cf.

$$\left\langle \overline{q}(0)q(0)\right\rangle = \frac{1}{Z}\int dAe^{-S_{Glue}} \det\left[\mathcal{D}+m\right] \operatorname{Tr}\left[\left(0\left|\frac{-1}{\mathcal{D}+m}\right|0\right)\right] = \left\langle \pi\rho\left(\lambda=0\right)\right\rangle$$

$$\rightarrow i\mathcal{D}\psi_{\lambda} = \lambda\psi_{\lambda} \quad \text{where} \quad \psi_{\lambda}\left(0\right) = \left(0\left|\lambda\right) \qquad \rho(\lambda) = \frac{1}{V}\int d^{3}x\psi_{\lambda}^{+}(x)\psi_{\lambda}(x)$$

(SHL, S.Cho, IJMPE arXiv:1302.0642)

Chiral order parameters: V - A correlator + more

$$\square \nabla = \Pi^{VV} - \Pi^{AA} = \frac{1}{V} \int d^4 x \left[\left\langle \overline{q} \gamma^{\mu} \tau^a q(x), \overline{q} \gamma^{\mu} \tau^a q(0) \right\rangle - \left\langle \overline{q} \tau^a i \gamma^5 \gamma^{\mu} q(x), \overline{q} \tau^a i \gamma^5 \gamma^{\mu} q(0) \right\rangle \right]$$
$$= -\frac{1}{2} \operatorname{Tr} \left[\gamma^{\mu} \left(S(x,0) - i \gamma^5 S(x,0) i \gamma^5 \right) \gamma^{\mu} \left(S(0,x) - i \gamma^5 S(0,x) i \gamma^5 \right) \right] \propto \left\langle \rho^2 \left(\lambda = 0 \right) \right\rangle$$



Weinberg sum rule V

 $m_{\rho} = m_{a_1} = m_0$ when chiral symmetry is restored. What about m_0 ?

Vector meson mass in the chiral symmetry restored vacuum

• QCD sum rule for ρ and a_1 meson

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Jisu Kim, SHL:PRD103(2021)L051501

$$\Pi^{VV} = \dots \frac{1}{Q^{6}} \left[-2\pi\alpha \left\langle \left(\bar{q}\gamma_{\mu}\gamma^{5}\lambda^{a}\tau^{3}q \right)^{2} \right\rangle - \frac{4\pi\alpha}{9} \left\langle \left(\sum_{ud} \bar{q}\gamma_{\mu}\lambda^{a}q \right) \left(\sum_{uds} \bar{q}\gamma_{\mu}\lambda^{a}q \right) \right\rangle \right]$$
$$\Pi^{AA} = \dots \frac{1}{Q^{6}} \left[-2\pi\alpha \left\langle \left(\bar{q}\gamma_{\mu}\lambda^{a}\tau^{3}q \right)^{2} \right\rangle - \frac{4\pi\alpha}{9} \left\langle \left(\sum_{ud} \bar{q}\gamma_{\mu}\lambda^{a}q \right) \left(\sum_{uds} \bar{q}\gamma_{\mu}\lambda^{a}q \right) \right\rangle \right]$$

$$\left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle = \frac{1}{2} \left[\left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} + \left(\overline{q} \gamma_{\mu} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle \right]_{S} + \frac{1}{2} \left[\left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} - \left(\overline{q} \gamma_{\mu} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle \right]_{B} \\ \left\langle \left(\overline{q} \gamma_{\mu} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle = \frac{1}{2} \left[\left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} + \left(\overline{q} \gamma_{\mu} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle \right]_{S} - \frac{1}{2} \left[\left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} - \left(\overline{q} \gamma_{\mu} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle \right]_{B} \\ no \text{ contribution from } \rho(\lambda = 0) \qquad \pm \qquad \left\langle \left(\rho(\lambda = 0) \right)^{2} \right\rangle \propto \left\langle \overline{q} q \right\rangle^{2}$$

$$\Pi^{VV} = \dots \frac{1}{Q^{6}} \left[\frac{14}{9} \langle B \rangle + \langle S \rangle \right] , \qquad \Pi^{AA} = \dots \frac{1}{Q^{6}} \left[-\frac{22}{9} \langle B \rangle + \langle S \rangle \right]$$
$$\langle B \rangle = -\pi \alpha \left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle_{B} \xrightarrow{\text{Vacuum Saturation}} = -\pi \alpha \frac{8}{9} \langle \overline{q} q \rangle^{2} \qquad \times \kappa \text{ in previous sum rules}$$
$$\langle S \rangle = -\frac{22\pi \alpha}{9} \left\langle \left(\overline{q} \gamma_{\mu} \gamma^{5} \lambda^{a} \tau^{3} q \right)^{2} \right\rangle_{S} + \dots \xrightarrow{\text{Vacuum Saturation}} = 0$$

 $\langle B \rangle$ and $\langle S \rangle$ can be determined separately from ρ and a_1 sum rules



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Pole	$B(\text{GeV}^6)$	$S(\text{GeV}^6)$
δ	7.42×10^{-4}	5.65×10^{-4}
$\operatorname{B-W}^1$	6.42×10^{-4}	6.05×10^{-4}
$\mathrm{B}\text{-}\mathrm{W}^2$	5.75×10^{-4}	7.11×10^{-4}

$$\Pi^{VV} = \dots \frac{1}{Q^6} \left[\frac{14}{9} \langle B \rangle + \langle S \rangle \right] , \qquad \Pi^{AA} = \dots \frac{1}{Q^6} \left[-\frac{22}{9} \langle B \rangle + \langle S \rangle \right]$$

Keep $\langle S \rangle$ fixed and change $\langle B \rangle$ to $0.7 \times \langle B \rangle$ and $0 \times \langle B \rangle \rightarrow \langle \overline{q}q \rangle = 0$



 $m_{\rho} = m_{a_1} = m_0 \sim 550 \pm 50$ MeV in the chiral symmetry restored vacuum $\Delta m_{\rho} \sim -100$ MeV from purely partial chiral symmetry restoration in nuclear matter

- 1. Confinement, chiral symmetry breaking, $U_A(1)$ effects all have different origin and contribute to hadron mass
- 2. Mass difference between chiral partners are directly related to chiral symmetry breaking $\langle VV AA \rangle$
- 3. Since the origin of chiral symmetry breaking effects is identified, Dividing quark operators into chiral symmetric and breaking operators in QCD sum rules, one can relate individual mass to chiral symmetry restoration $\langle VV \rangle$, $\langle AA \rangle$

How can we observe mass shift – small width hadrons

KEK E325, J-PARC E16



Vacuum values	Mass	Width
φ	1020 MeV	4.266 MeV

How can we observe mass shift – small width hadrons

CBELSA/TAPS coll (V. Metag, M. Nanova et al)



Vacuum values	Mass	Width
ω	782.65 MeV	8.49 MeV
η'	957.78 MeV	0.198 MeV

Mass shift by V. Metag (PPNP97 (2017)199)

Downward mass shift at nuclear matter



Width increase at nuclear matter

Lesson from experiment

- 1. Look at small width hadrons (<100 MeV)
- 2. Can look at excitation energy \rightarrow mass shift
- 3. Look at transparency \rightarrow Width

Small vacuum width and chiral partner

1. $f_1(1285)$ and ω

2. K* and K₁

Light vector mesons – chiral partners ?

12	J ^{PC} =1	Mass	Width	J ^{PC} =1 ⁺⁺	Mass	Width
	ρ	770	150.	a ₁	1260	250-600
	ω	782	8.49	f ₁	1285	24.2
	φ	1020	4.266	f ₁	1420	54.9
	K*(1⁻)	892	50.3	K ₁ (1+)	1270	90

 \checkmark (ρ , a_1) are chiral partners but have large vacuum width

$$\rho \rightarrow \left(\overline{q}_{R} \gamma_{\mu} \tau q_{R} + \overline{q}_{L} \gamma_{\mu} \tau q_{L}\right) \qquad a_{1} \rightarrow \left(\overline{q}_{R} \gamma_{\mu} \tau q_{R} - \overline{q}_{L} \gamma_{\mu} \tau q_{L}\right)$$

Light vector mesons – chiral partners ?

J ^{PC} =1	Mass	Width	J ^{PC} =1 ⁺⁺	Mass	Width
ρ	770	150.	a₁	1260	250-600
ω	782	8.49	f ₁	1285	24.2
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K*(1⁻)	892	50.3	K ₁ (1+)	1270	90

Coupling to quark currents [Gubler, Kunihiro, Lee, PLB767(2017)336]

$$\begin{split} \omega \to & \left(\overline{u} \gamma_{\mu} u + \overline{d} \gamma_{\mu} d \right) & f_1 (1285) \to \left(\overline{u} \gamma_{\mu} \gamma^5 u + \overline{d} \gamma_{\mu} \gamma^5 d \right) \\ \phi \to & \left(\overline{s} \gamma_{\mu} s \right) & f_1 (1420) \to \left(\overline{s} \gamma_{\mu} \gamma^5 s \right) \end{split}$$

But they are not chiral partners \rightarrow Jisu Kim and SHL (in preparation)

$$\Pi^{\omega\omega} - \Pi^{f_1 f_1} = \left\langle \left(\overline{u}_R \gamma_\mu u_R \right) \left(\overline{u}_L \gamma_\mu u_L \right) + \left(\overline{u}_R \gamma_\mu u_R \right) \left(\overline{d}_L \gamma_\mu d_L \right) \right\rangle$$
$$= \Pi^{\rho\rho} - \Pi^{a_1 a_1} + \left\langle 2 \left(\overline{u}_R \gamma_\mu u_R \right) \left(\overline{d}_L \gamma_\mu d_L \right) \right\rangle$$

 $\langle \overline{q}q \rangle$ Chiral symmetric

J ^{PC} =1	Mass	Width	J ^{PC} =1 ⁺⁺	Mass	Width
ρ	770	150.	a ₁	1260	250-600
ω	782	8.49	f ₁	1285	24.2
φ	1020	4.266	f ₁	1420	54.9
K*(1⁻)	892	50.3	K ₁ (1 ⁺)	1270	90

 (ρ, a_1) are chiral partners but have too large vacuum width

$$\rho \to \left(\overline{q}_R \gamma_\mu \tau q_R + \overline{q}_L \gamma_\mu \tau q_L\right) \qquad a_1 \to \left(\overline{q}_R \gamma_\mu \tau q_R - \overline{q}_L \gamma_\mu \tau q_L\right)$$

Coupling to quark currents

$$\omega \to \left(\overline{u} \gamma_{\mu} u + \overline{d} \gamma_{\mu} d\right) \quad \phi \to \left(\overline{s} \gamma_{\mu} s\right) \qquad K^* \to \left(\overline{q} \gamma_{\mu} s\right), \quad \left(\overline{s} \gamma_{\mu} q\right)$$

 \rightarrow What about quark content of K₁?

Solution Are (K^*, K_1) chiral partners ?

K₁(1270) sum rules in medium (Song, Hatsuda, Lee, PLB792 (2019) 160)



$$K^* \to \left(\overline{u} \gamma_{\mu} s \right) \qquad \qquad K_1 (1270) \to \left(\overline{u} \gamma_{\mu} \gamma^5 s \right)$$

Chiral Partner ?

Yes, they are chiral partners

$$\Pi^{\rho\rho} - \Pi^{a_{1}a_{1}} = \left\langle \left(\overline{u}_{R} \gamma_{\mu} u_{R} \right) \left(\overline{u}_{L} \gamma_{\mu} u_{L} \right) - \left(\overline{u}_{R} \gamma_{\mu} u_{R} \right) \left(\overline{d}_{L} \gamma_{\mu} d_{L} \right) \right\rangle \propto \left\langle \overline{q}q \right\rangle^{2} \sim \left(m_{a_{1}} - m_{\rho} \right) \approx 490 \text{ MeV}$$

$$= -\frac{1}{2} \text{Tr} \left[\gamma^{\mu} \left(S_{q,s} \left(x, 0 \right) - i\gamma^{5} S_{q,s} \left(x, 0 \right) i\gamma^{5} \right) \gamma^{\mu} \left(S_{q} \left(0, x \right) - i\gamma^{5} S_{q} \left(0, x \right) i\gamma^{5} \right) \right]$$

$$\Pi^{K^{*}K^{*}} - \Pi^{K_{1}K_{1}} = \left\langle \left(\overline{u}_{R} \gamma_{\mu} s_{R} \right) \left(\overline{s}_{L} \gamma_{\mu} u_{L} \right) \right\rangle \qquad \propto \left\langle \overline{q}q \right\rangle \left\langle \overline{s}s \right\rangle \sim \left(m_{K_{1}} - m_{K^{*}} \right) \approx 378 \text{ MeV}$$

The Distinct spectral density \rightarrow can understand how chiral symmetry restoration is realized in nature



• Expected mass shift from sum rules



Hence, mass shift at nuclear matter $\Delta m(K_1^-) \approx -208 \text{ MeV} \qquad \Delta m(K_1^+) \approx +32 \text{ MeV}$ Possible future experiment

 \rightarrow K₁ excitation energy measurement at JPARC



K₁/K* enhancement in Heavy Ion collision [H. Sung, Cho, Hong, Lee S. Lim PLB 819 (2021) 136388]

- Chemical freeze-out temperature in heavy ion collision at LHC: 156 MeV
- Chiral order parameter at 156 MeV
 substantially reduced

[Ding et al. arXiv:1312.0119]



Number of K1 and K* will be similar at the chemical freeze-out point after heavy ion collision



	Centrality dependence of hadron phase life time	Centrality (%) 0 - 5% 40 - 50% 70 - 80%	T _f (MeV) 90 108 147	t _c (fm/c) 8.7 4.9 2.2	t _f (fm/c) 28.1 13 2.9
¢	$0 \sim 5\% \text{ (centrality)}$ $t_c \sim 8.7 (fm/c) \qquad t_f$	 28.1(••••••••••••••••••••••••••••••••••••••)	
Ş	$70 \sim 80\% \text{ (centrality)}$ $70 \sim 80\% \text{ (centrality)}$ $t_{c} \sim 2.2(fm/c) \rightarrow t_{f} \sim 2.9(fm/c)$)			24

Centrality dependence of hadron phase life time

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Centrality dependence of hadron phase life time

Centrality (%)	T_f (MeV)	t _c (fm/c)	t _f (fm/c)
0 - 5%	90	8.7	28.1
40 - 50%	108	4.9	13
70 - 80%	147	2.2	2.9

Centrality dependence of final hadron yield ratio



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- 1. Mass difference between chiral partners are directly related to chiral symmetry breaking ρa_1 , $K^* K_1$ (small width)
- 2. Still, separating the 4-quark operators into chiral symmetric and breaking operators in QCD sum rules, one can identify the mass in the chiral symmetry restored vacuum. ρ , a_1 , K^* , K_1 , ϕ , f_1 ,.....
- 3. Experimental observation of mass shift of above or other particle in JPARC would be crucial : Looking forward to results on ϕ E-16
- 4. K^* , K_1 mass shift in nuclear matter can be done in JPARC
- 5. K_1/K^* measurement in heavy ion collision could be signature of chiral symmetry restoration in heavy ion collision